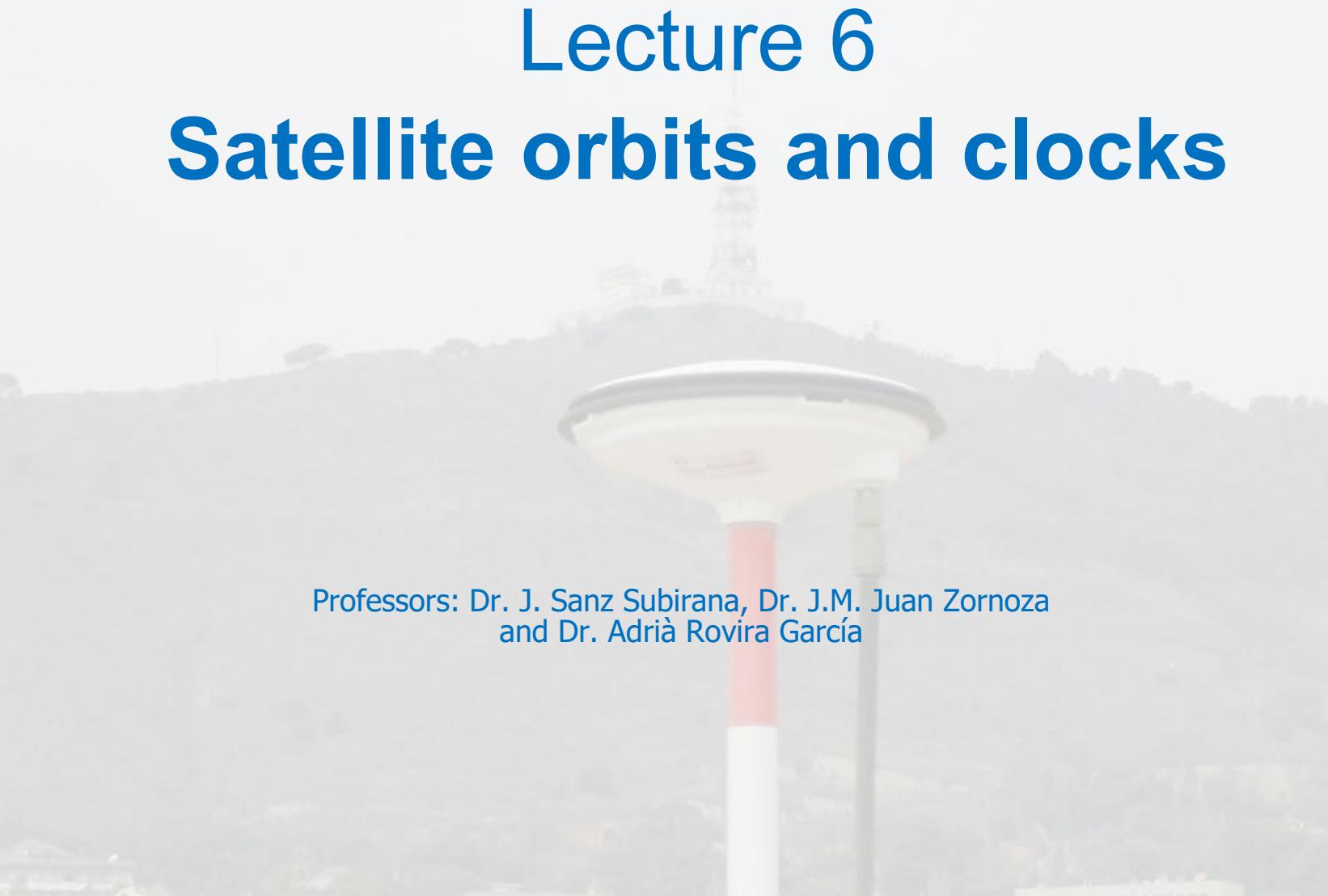


# Lecture 6

# Satellite orbits and clocks



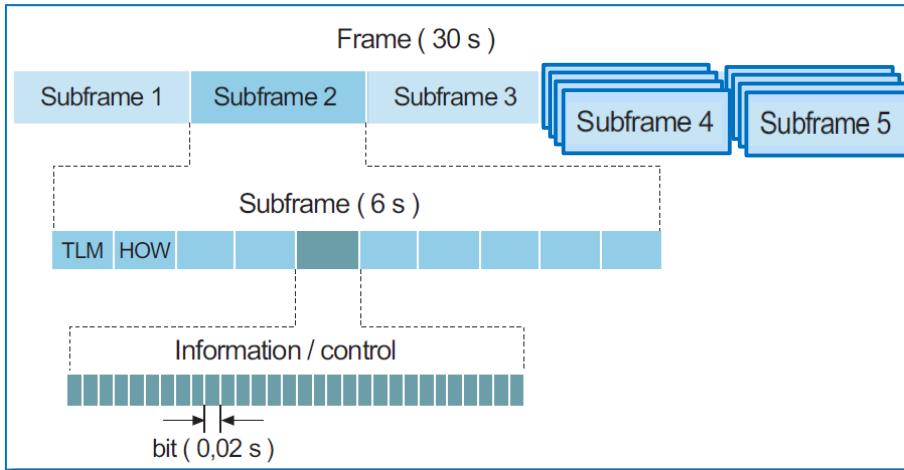
Professors: Dr. J. Sanz Subirana, Dr. J.M. Juan Zornoza  
and Dr. Adrià Rovira García

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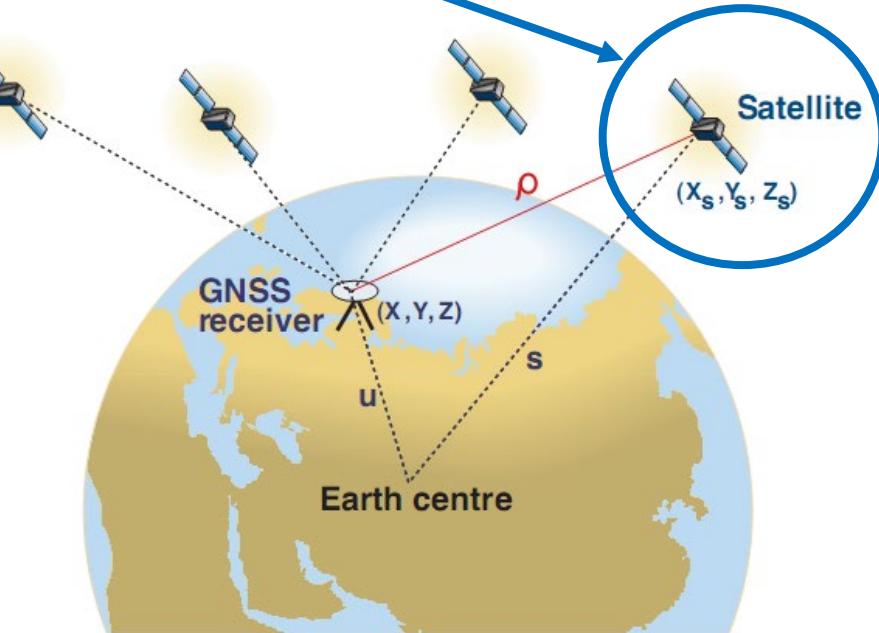
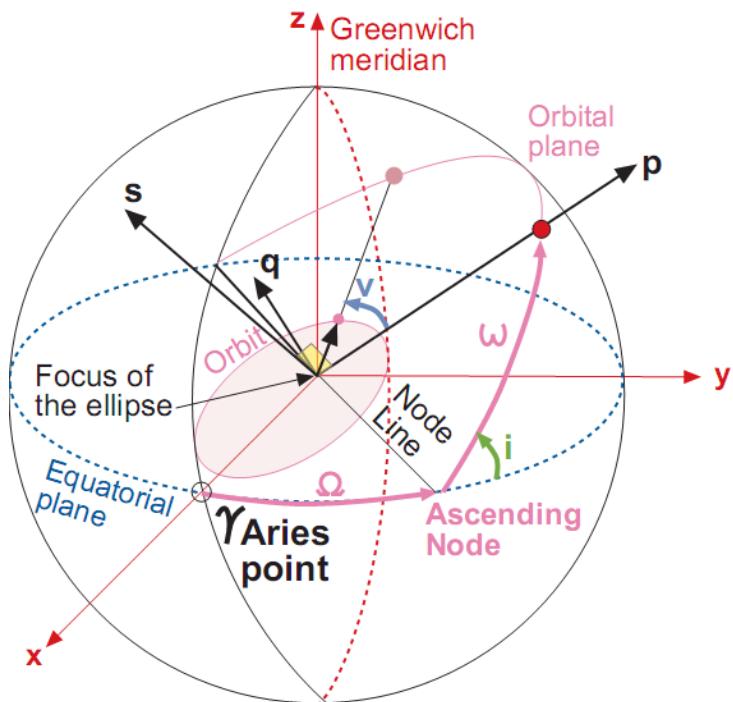
1. Elliptic orbit: Keplerian elements.
2. Perturbed Keplerian orbits: Osculating orbit.
3. GPS satellite coordinates computation and accuracy
  - 3.1. From Broadcast Navigation Message.
  - 3.2. From precise products.
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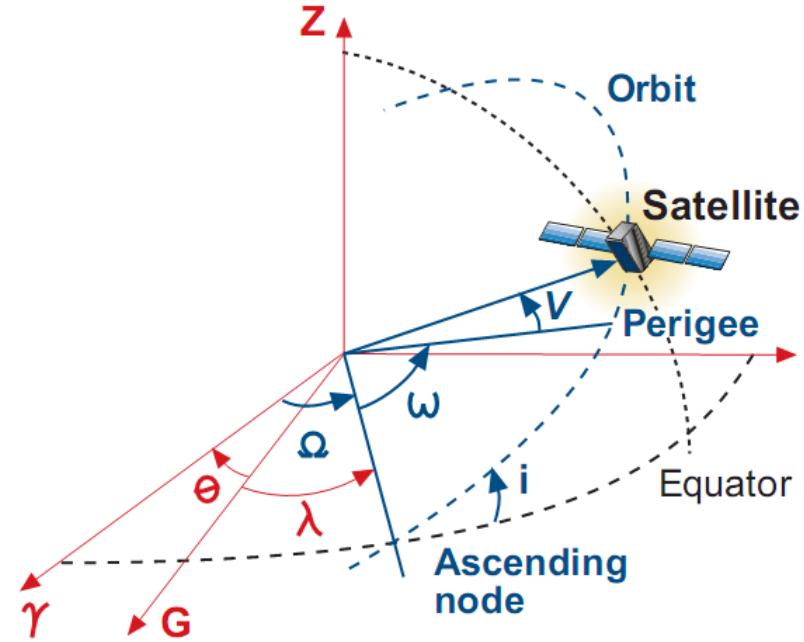
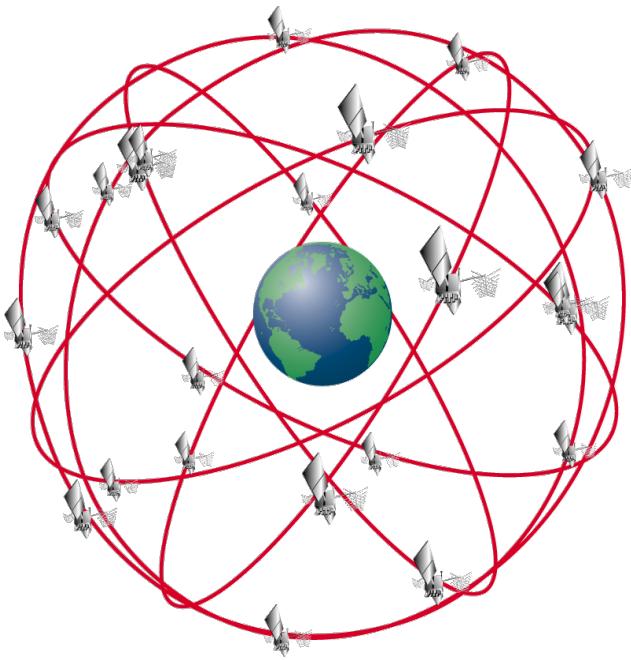


The GPS navigation message provides pseudo-Keplerian elements to compute satellite coordinates



$$(X, Y, Z, Vx, Vy, Vz) \rightarrow (a, e, i, \Omega, \omega, V)$$

**6 values** are needed ( $x, y, z, vx, vy, vz$ ) to provide the position and velocity of a body. They can be mapped into the **six Keplerian elements** ( $a, e, i, \Omega, \omega, V$ ), which provides the "natural" representation of the orbit!

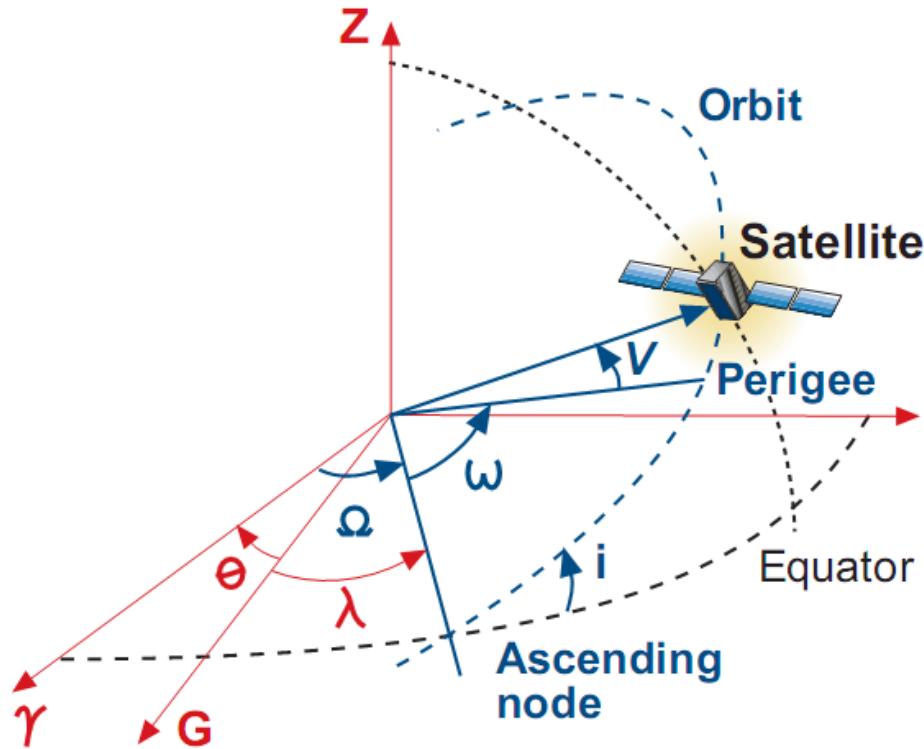
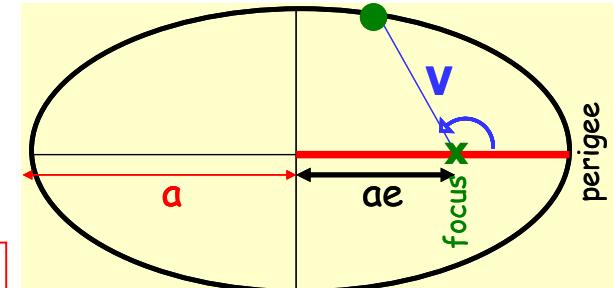


( $a, e,$   $i, \Omega, \omega,$   $\nu$ )

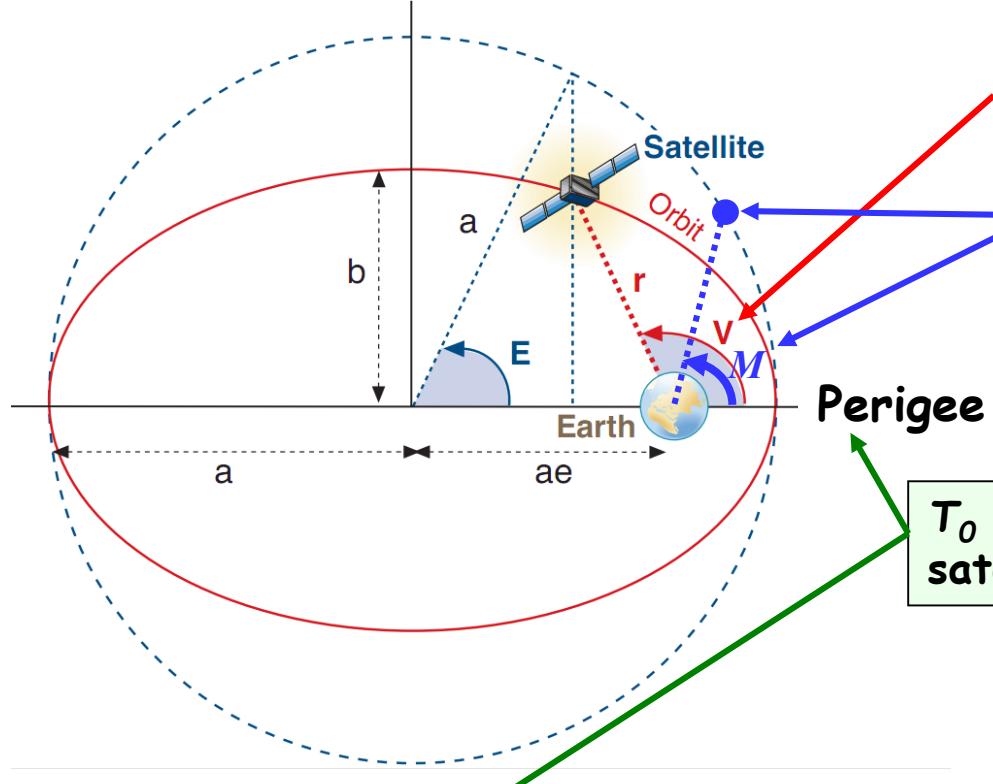
orbit  
shape

orbit  
orientation

position in  
the orbit



- $i$  inclination
- $\omega$  argument of perigee
- $\Omega$  arg. ascending node (Aries)
- $\lambda$  arg. ascending node (Greenwich)
- $\nu$  true anomaly
- $\theta$  sidereal time
- $\gamma$  vernal equinox
- $G$  Greenwich meridian



True anomaly  $V(t)$

Fictitious body moving at velocity  $n=2\pi/P=\text{constant}$   
 → Mean anomaly  $M(t)$

$T_0$  : time of passage by satellite's perigee

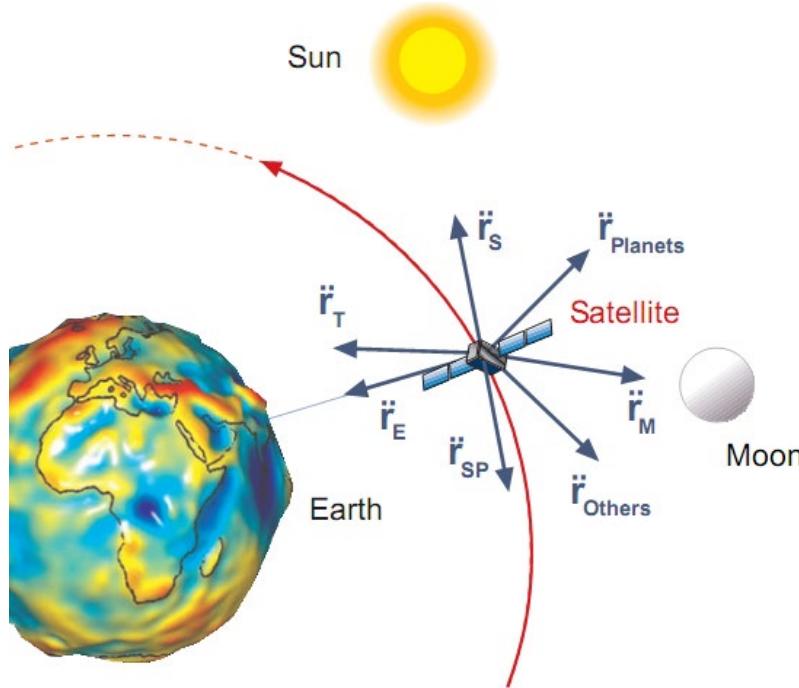
$$t \rightarrow \boxed{n = \frac{2\pi}{P} \quad T_0 \quad V(t)} \quad \mu, a, e$$

$$\begin{aligned} M(t) &= n(t - T_0) \quad ; \quad n = \frac{2\pi}{P} = \sqrt{\frac{\mu}{a^3}} \\ E(t) &= M(t) + e \sin E(t) \\ V(t) &= 2 \arctan \left[ \sqrt{\frac{1+e}{1-e}} \tan \frac{E(t)}{2} \right] \end{aligned}$$

# Contents

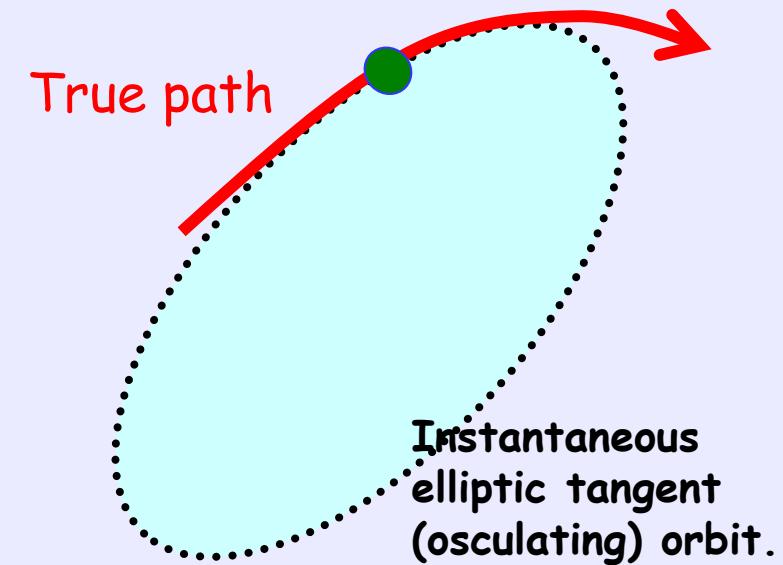
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Due to the non-spherical nature of gravitational potential, the attraction of the Sun and Moon, the solar radiation pressure, etc., **the true satellite path deviates from the elliptic orbit.**



At any time an elliptical orbit tangent to the true path can be defined. This is the “osculating orbit”, whose Keplerian elements vary with time “t”:

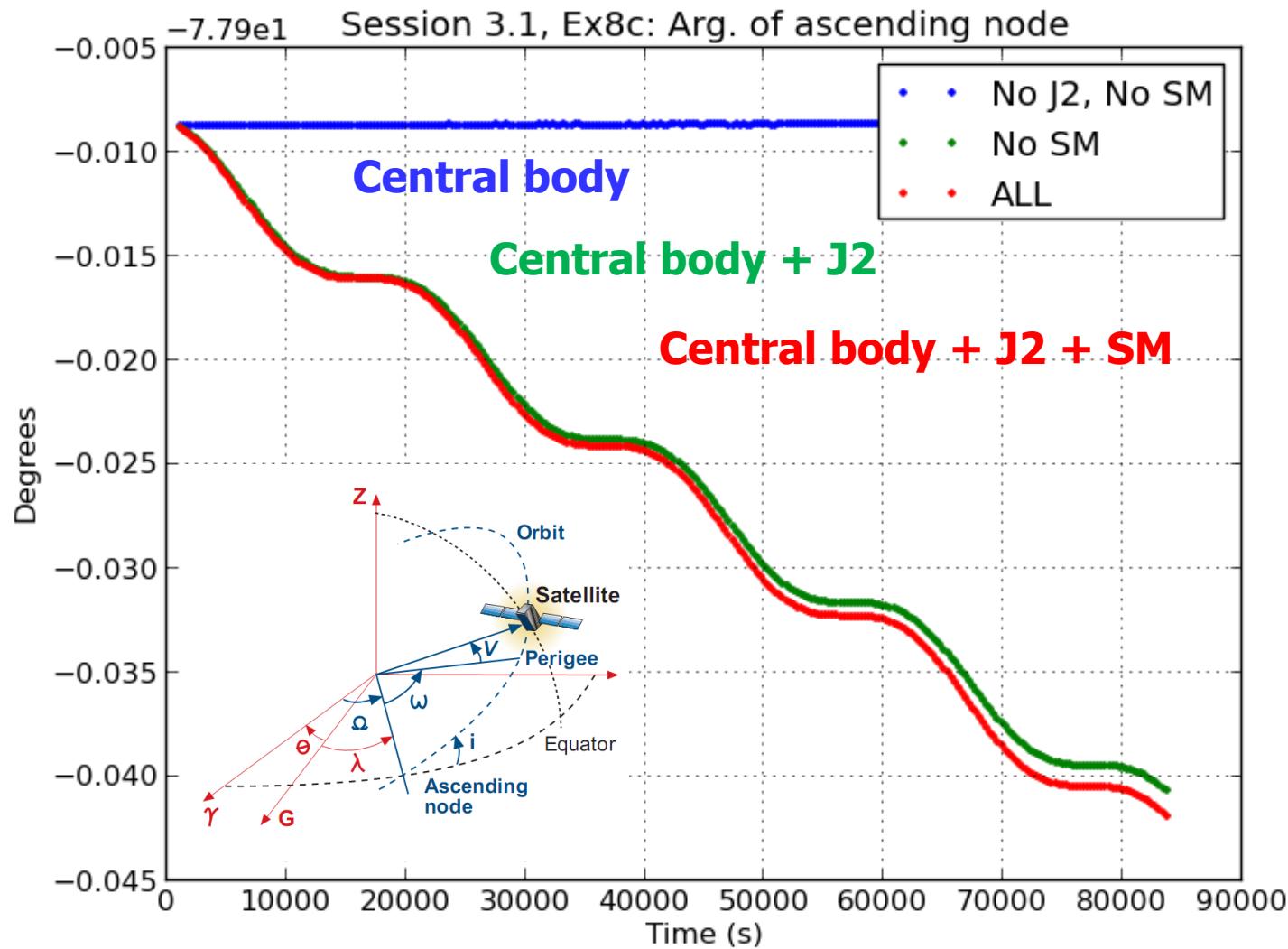
$$a(t), e(t), i(t), \Omega(t), \omega(t), V(t)$$

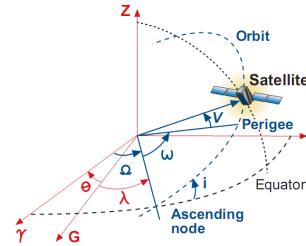
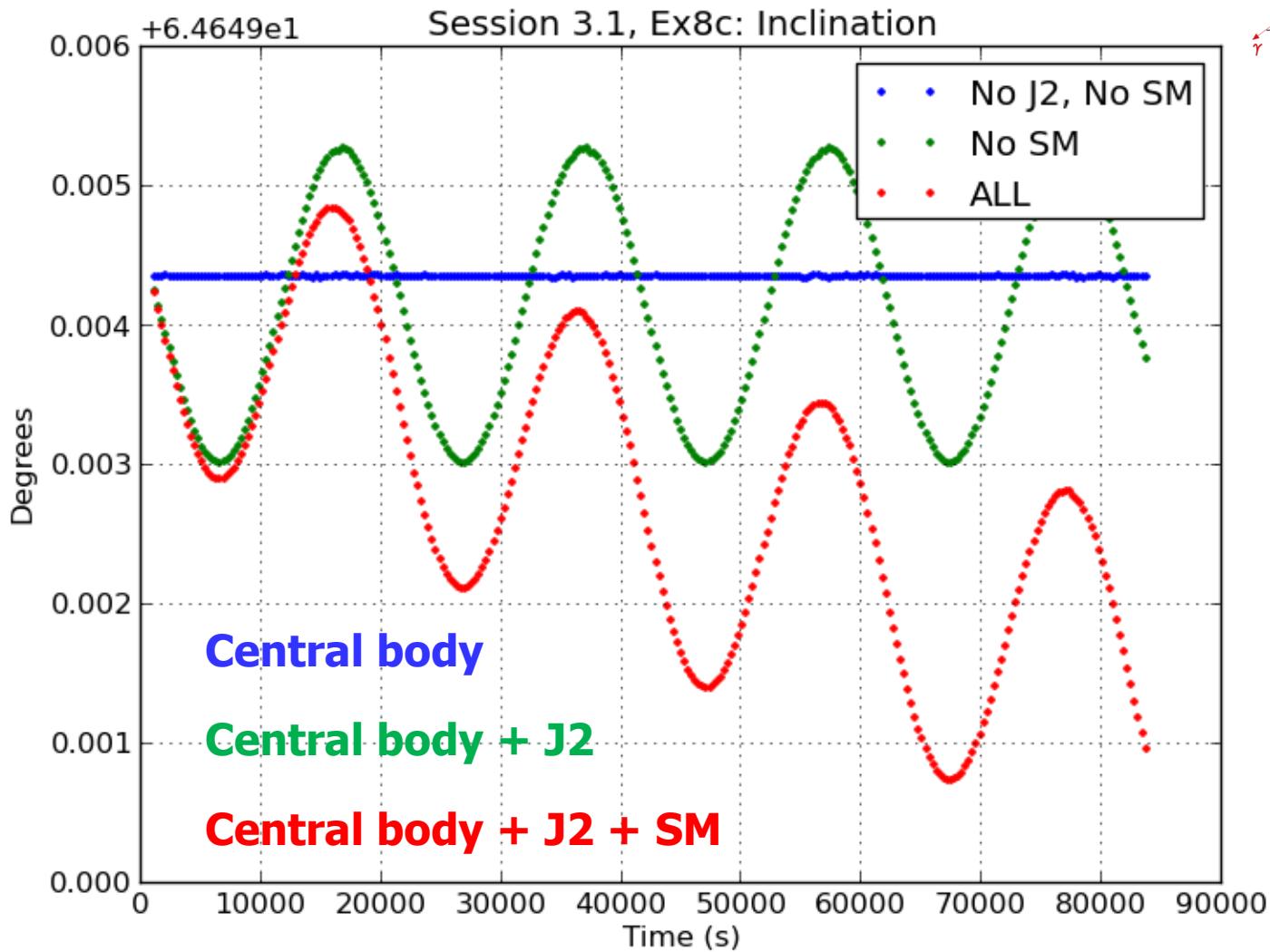


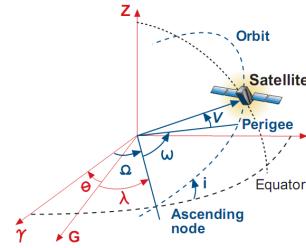
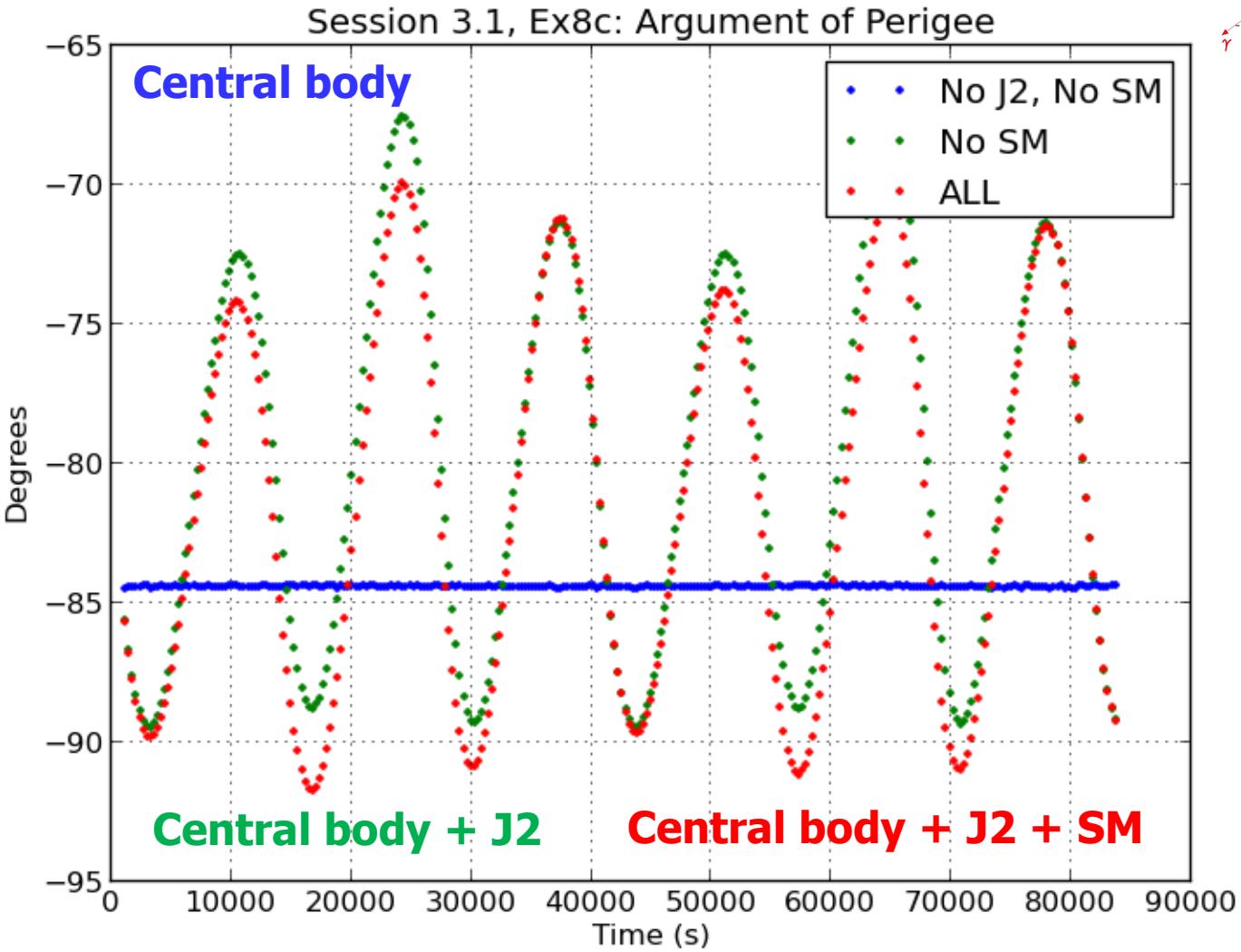
# Different magnitudes of perturbation and their effects on GPS orbits

Perturbation	Acceleration (m/s <sup>2</sup> )	Orbital effect	
		in 3 hours	in 3 days
Central force (as a reference)	0.56		
$J_2$	$5 \cdot 10^{-5}$	2 km	14 km
Rest of the harmonics	$3 \cdot 10^{-7}$	50–80 m	100–1500 m
Solar + Moon grav.	$5 \cdot 10^{-6}$	5–150 m	1000–3000 m
Tidal effects	$1 \cdot 10^{-9}$	–	0.5–1.0 m
Solar rad. pressure	$1 \cdot 10^{-7}$	5–10 m	100–800 m

GLONASS Broadcast orbit integration terms







## Calculation of osculating orbital elements from position and velocity (**rv2osc.f**)

$$(x, y, z, v_x, v_z, v_z) \Rightarrow (a, e, i, \Omega, \omega, M)$$

$$\vec{c} = \vec{r} \times \vec{v} \implies p = \frac{c^2}{\mu} \implies p$$

$$v^2 = \mu(2/r - 1/a) \implies a$$

$$p = a(1 - e^2) \implies e$$

$$\vec{c} = c \vec{S} \implies \Omega = \arctan(-c_x/c_y); \quad i = \arccos(c_z/c) \implies \Omega, i$$

$$\begin{pmatrix} x \\ y \\ z \end{pmatrix} = R \begin{pmatrix} r \cos(V) \\ r \sin(V) \\ 0 \end{pmatrix} = r \begin{pmatrix} \cos \Omega \cos(\omega + V) - \sin \Omega \sin(\omega + V) \cos i \\ \sin \Omega \cos(\omega + V) + \cos \Omega \sin(\omega + V) \cos i \\ \sin(\omega + V) \sin i \end{pmatrix} \implies \omega + V$$

$$r = \frac{p}{1 + e \cos(V)} \implies \omega, V$$

$$\tan(E/2) = \left(\frac{1-e}{1+e}\right)^{1/2} \tan(V/2) \quad ; \quad M = E - e \sin E \implies M$$

# Calculation of position and velocity from osculating orbital elements (osc2rv.f)

$$(a, e, i, \Omega, \omega, \underbrace{T; t}_V) \Rightarrow (x, y, z, v_x, v_y, v_z)$$

$t \Rightarrow M$	$M \Rightarrow E$	$E \Rightarrow (r, V)$
$M = n(t - T)$	$M = E - e \sin E$	$r = a(1 - e \cos E)$
		$\tan(V/2) = (\frac{1+e}{1-e})^{1/2} \tan(E/2)$

$$\begin{pmatrix} x \\ y \\ z \end{pmatrix} = R \begin{pmatrix} r \cos(V) \\ r \sin(V) \\ 0 \end{pmatrix} ; \begin{pmatrix} v_x \\ v_y \\ v_z \end{pmatrix} = \frac{na^2}{r} \{ \vec{Q}(1 - e^2)^{1/2} \cos E - \vec{P} \sin E \}$$

**Where:**

$$\begin{aligned}
 R &= R_3(-\Omega)R_1(-i)R_3(-\omega) = \\
 &= \begin{pmatrix} \cos \Omega & -\sin \Omega & 0 \\ \sin \Omega & \cos \Omega & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos i & -\sin i \\ 0 & \sin i & \cos i \end{pmatrix} \begin{pmatrix} \cos \omega & -\sin \omega & 0 \\ \sin \omega & \cos \omega & 0 \\ 0 & 0 & 1 \end{pmatrix} \\
 &= \begin{pmatrix} P_x & Q_x & S_x \\ P_y & Q_y & S_y \\ P_z & Q_z & S_z \end{pmatrix} = [\vec{P} \ \vec{Q} \ \vec{S}]
 \end{aligned}$$

# Exercise: Orbital elements variation:

File 1995-10-18.eci contains the precise position and velocities of GPS satellites every 5 minutes for October 18th, 1995 in a Earth-Centred Inertial system (ECI)

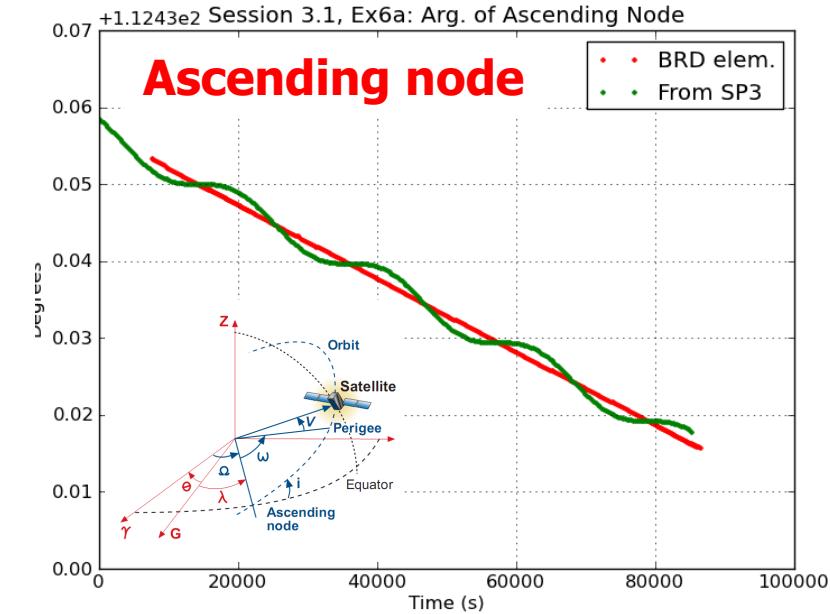
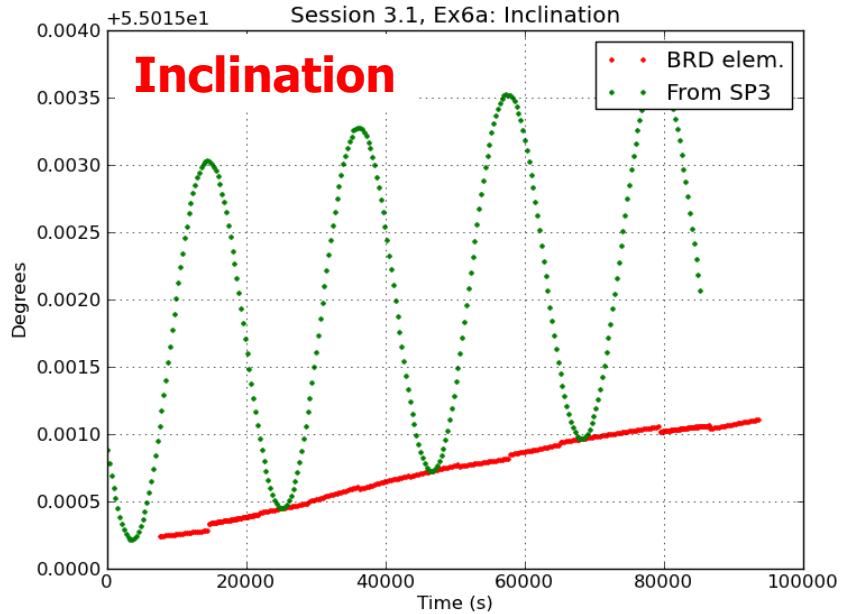
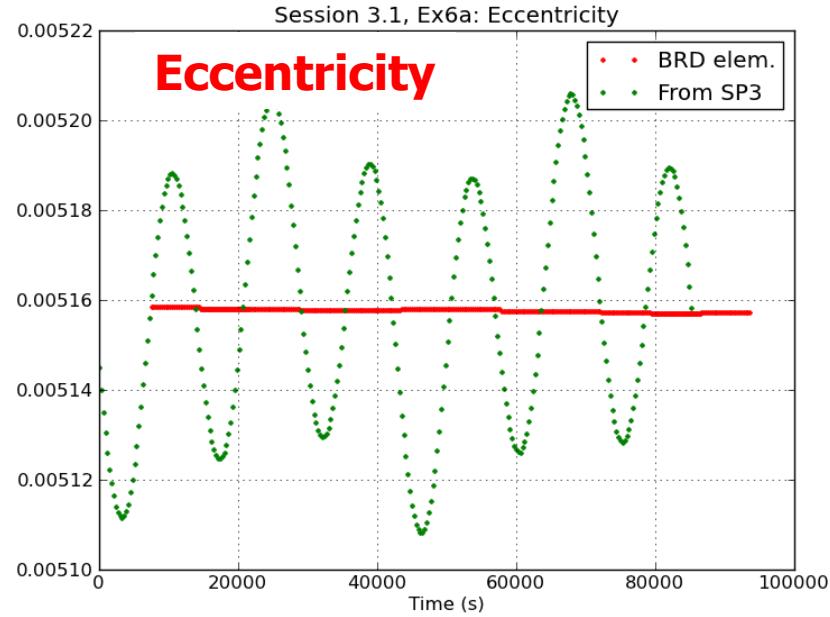
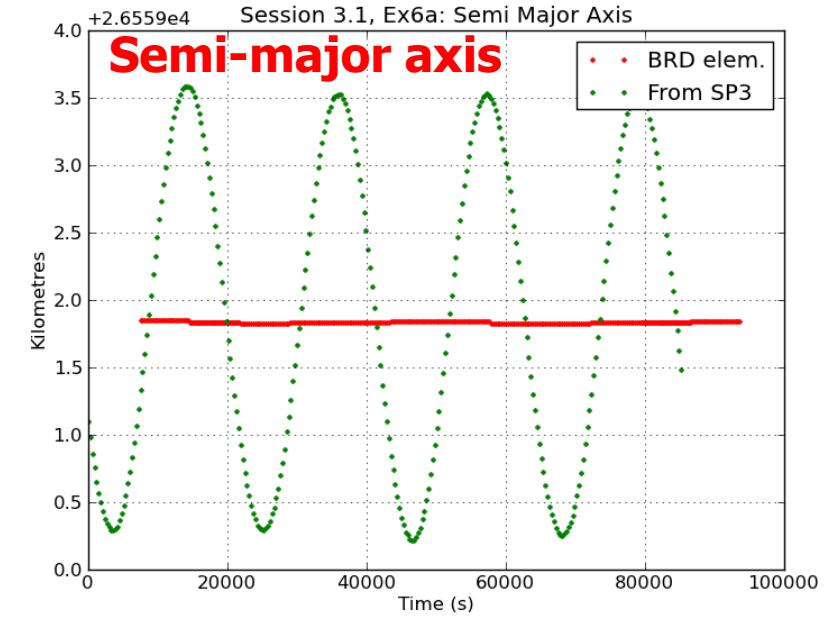
[from JPL/NASA server:

[ftp://sideshow.jpl.nasa.gov/pub/gipsy\\_products](ftp://sideshow.jpl.nasa.gov/pub/gipsy_products)

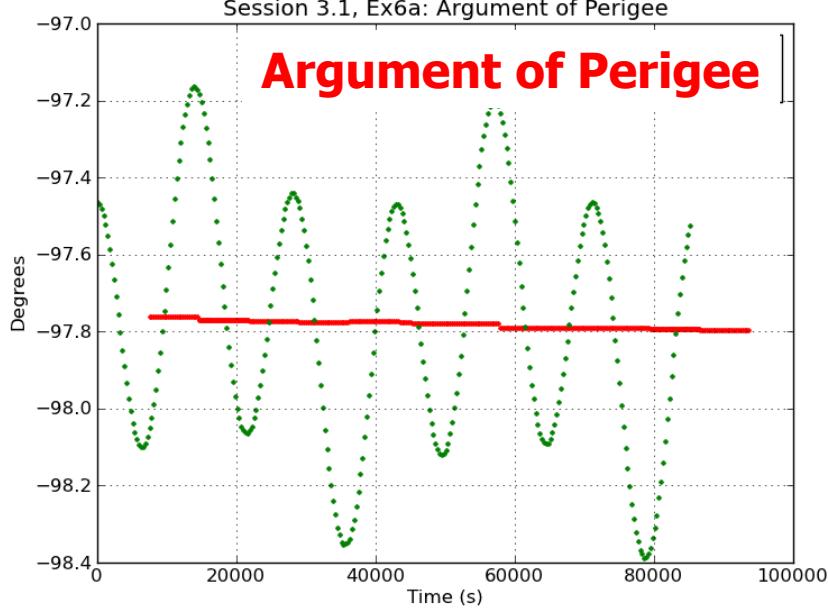
- a) Use program “**rv2osc**” to compute the instantaneous orbital elements  $(X, Y, Z, V_x, V_y, V_z) \rightarrow (a, e, i, \Omega, \omega, V)$
- b) Plot the orbital elements in function of time to show their variation:  $a(t), e(t), i(t), \Omega(t), \omega(t), V(t)$
- c) Compare with the broadcast orbital elements

## Solution:

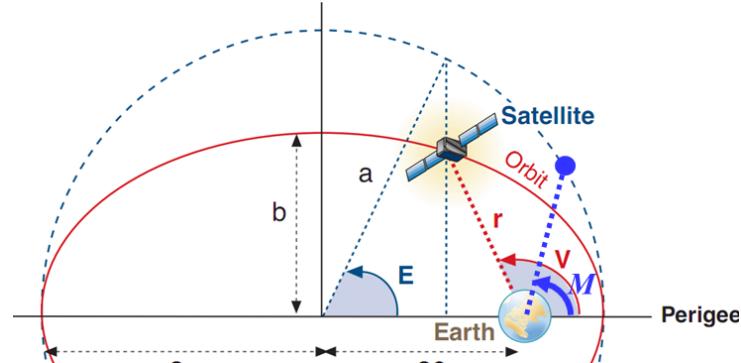
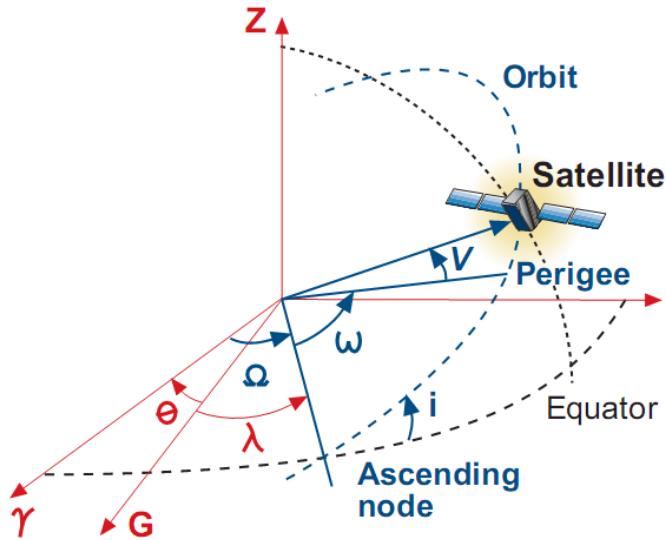
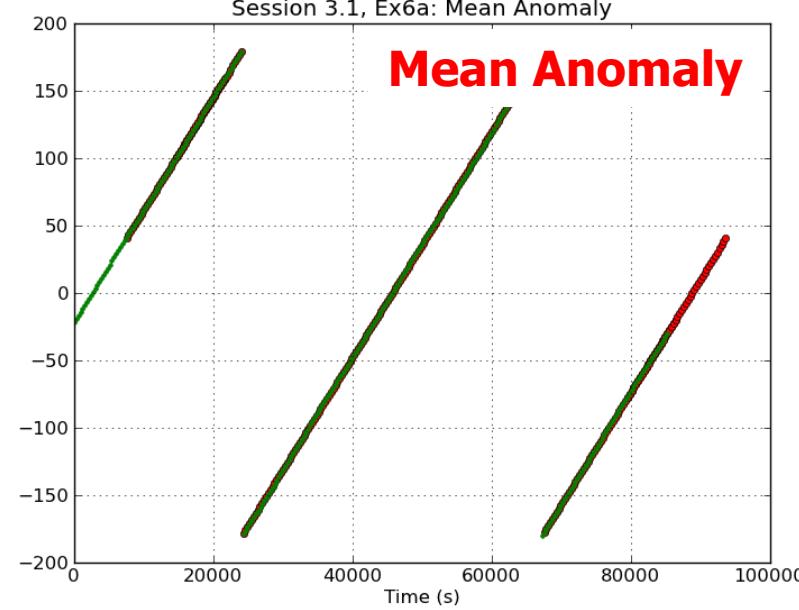
- a) cat 1995-10-18.eci|rv2osc> orb.dat
- b) See the following plots



Session 3.1, Ex6a: Argument of Perigee



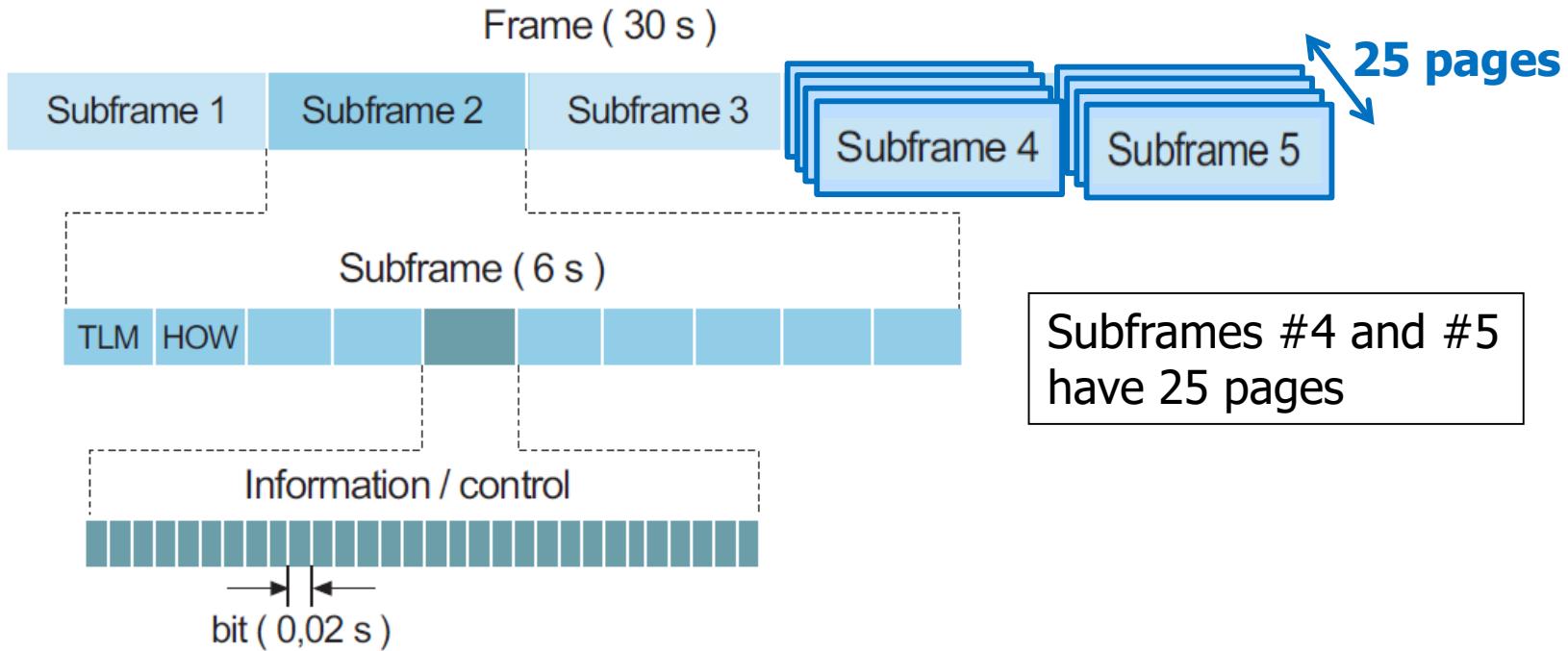
Session 3.1, Ex6a: Mean Anomaly



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# GPS navigation message



**One Master Frame** includes All 25 pages of Subframes #4 and #5 →  $25 \times 30\text{s} = 12.5 \text{ min}$

**Subframe 1** contains information about the parameters to be applied to **satellite clock** status for its correction. These values are polynomial coefficients that allow time onboard to be converted to GPS time. The subframe also contains information on satellite health condition.

**Subframes 2 and 3** contain **satellite ephemerides**.

**Subframe 4** provides **ionospheric model** parameters (in order to adjust for ionospheric refraction), UTC information, part of the **almanac**, and indications whether the A/S is activated or not (which transforms the P code into encrypted Y code).

**Subframe 5** contains data from the **almanac** and on constellation status. It allows rapid identification of the satellite from which the signal comes. A total of 25 frames are needed to complete the almanac.

# Ephemeris in navigation message

Parameter	Explanation
$IODE$	Series number of ephemerides data
$t_{oe}$	Ephemerides reference epoch
$\sqrt{a}$	Square root of semi-major axis
$e$	Eccentricity
$M_o$	Mean anomaly at reference epoch
$\omega$	Argument of perigee
$i_o$	Inclination at reference epoch
$\Omega$	Ascending node's right ascension
$\Delta n$	Mean motion difference
$\bullet$ $i$	rate of inclination angle
$\bullet$ $\Omega$	Rate of node's right ascension
$c_{uc}, c_{us}$	Latitude argument correction
$c_{rc}, c_{rs}$	Orbital radius correction
$c_{ic}, c_{is}$	Inclination correction

In order to calculate WGS84 satellite coordinates, you should apply the following algorithm [GPS/SPS-SS, table 2-15] (see in the book FORTRAN subroutine orbit.f)

# RINEX ephemeris file

2 NAVIGATION DATA GPS RINEX VERSION/ TYPE  
 XPRINT v1.1 gAGE 00/08/17 09:31:37 PGM / RUN BY / DATE  
 gAGE BROADCAST EPHEMERIS FILE COMMENT  
 $+1.7695E-08 +2.2352E-08 -1.1921E-07 -1.1921E-07$  ION ALPHA  
 $+1.1878E+05 +1.4746E+05 -1.3107E+05 -3.2768E+05$  ION BETA  
 $+1.955777406693E-08+1.598721155460E-14 405504$  1064 DELTA\_UTC: A0,A1,T,W  
 13 LEAP SECONDS  
 END OF HEADER

03 00 5 30 10 0 40.0	$+7.855705916882E-06+3.524291969370E-12+0.000000000000E+00$	
		Mo
$+1.010000000000E+02+6.500000000000E+01+5.456298524109E-09+5.530285585107E-01$		
$+3.475695848465E-06+1.308503560722E-03+2.641230821609E-06+5.153678266525E+03$		e, $\sqrt{a}$
$+2.088000000000E+05+1.117587089539E-08+7.472176136643E-01-1.862645149231E-09$		TOE, $\Omega$
$+9.412719852649E-01+3.163750000000E+02+1.125448382894E+00-8.826796182859E-09$		io, $\omega$
$+1.239337382719E-10+1.000000000000E+00+1.064000000000E+03+0.000000000000E+00$		
$+4.000000000000E+00+0.000000000000E+00-4.190951585770E-09+6.130000000000E+02$		TGD
$+2.044980000000E+05+0.000000000000E+00+0.000000000000E+00+0.000000000000E+00$		

06 00 5 30 10 0 0.0

$+1.636799424887E-06+0.000000000000E+00+0.000000000000E+00$	
$+6.000000000000E+01+5.100000000000E+01+5.198073527168E-09-5.601816471398E-01$	
$+2.635642886162E-06+6.763593177311E-03+2.468004822731E-06+5.153726325989E+03$	
$+2.088000000000E+05+1.862645149231E-08+7.894129138508E-01+8.195638656616E-08$	
$+9.487675576456E-01+3.229687500000E+02-2.409256713064E+00-8.734292400447E-09$	
$+4.714481929846E-11+1.000000000000E+00+1.064000000000E+03+0.000000000000E+00$	

### 3.1. Computation of satellite coordinates from navigation message (orbit.f)

- Computation of  $t_k$  time since ephemeris reference epoch  $t_{oe}$  ( $t$  and  $t_{oe}$  are given in GPS seconds of week):

$$t_k = t - t_{oe}$$

- Computation of mean anomaly  $M_k$  for  $t_k$ ,

$$M_k = M_0 + \left( \frac{\sqrt{\mu}}{\sqrt{a^3}} + \Delta n \right) t_k$$

- Iterative resolution of Kepler's equation in order to compute eccentric anomaly  $E_k$ :

$$M_k = E_k - e \sin E_k$$

- Calculation of true anomaly  $\nu_k$ :

$$\nu_k = \arctan \left( \frac{\sqrt{1-e^2} \sin E_k}{\cos E_k - e} \right)$$

- Computation of latitude argument  $u_k$  from perigee argument  $W$ , true anomaly  $\nu_k$  and corrections  $c_{uc}$  and  $c_{us}$ :

$$u_k = \omega + \nu_k + c_{uc} \cos 2(\omega + \nu_k) + c_{us} \sin 2(\omega + \nu_k)$$

- Computation of radial distance  $r_k$ , taking into consideration corrections  $c_{rc}$  and  $c_{rs}$ :

$$r_k = a(1 - 2 \cos E_k) + c_{rc} \cos 2(\omega + \nu_k) + c_{rs} \sin 2(\omega + \nu_k)$$

- Calculation of orbital plane inclination  $i_k$  from inclination  $i_0$  at reference epoch  $t_{oe}$  and corrections  $c_{ic}$  and  $c_{is}$ :

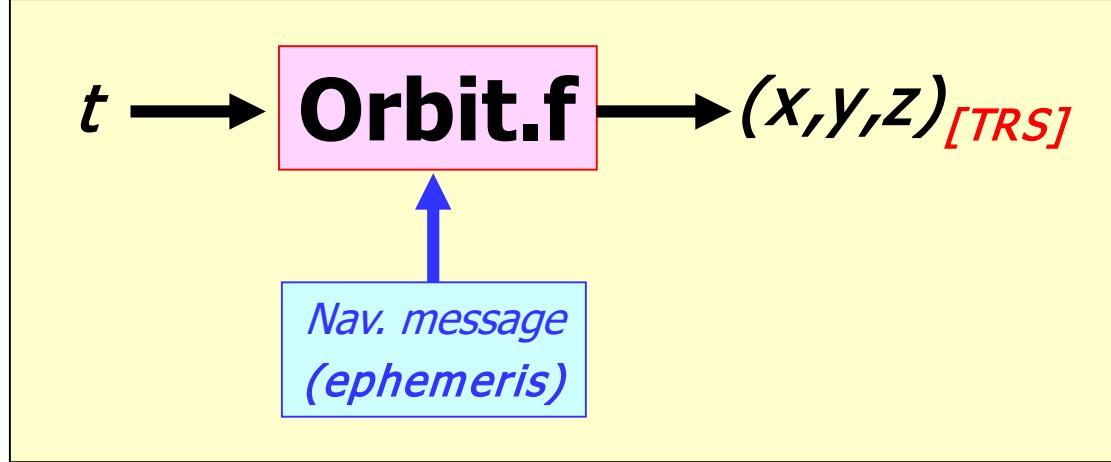
$$i_k = i_0 + it_k + c_{ic} \cos 2(\omega + \nu_k) + c_{is} \sin 2(\omega + \nu_k)$$

- Computation of ascending node longitude  $\Omega_k$  (Greenwich), from longitude  $\Omega_0$  at start of GPS week, corrected from apparent variation of sidereal time at Greenwich between start of week and reference time  $t_k = t - t_{oe}$ , and also corrected from change of ascending node longitude since reference epoch  $t_{oe}$ .

$$\Omega_k = \Omega_0 + (\Omega - \omega_E)t_k - \omega_E t_{oe}$$

- Calculation of coordinates in CTS system, applying three rotations (around  $u_k$ ,  $i_k$ ,  $\Omega_k$ ):

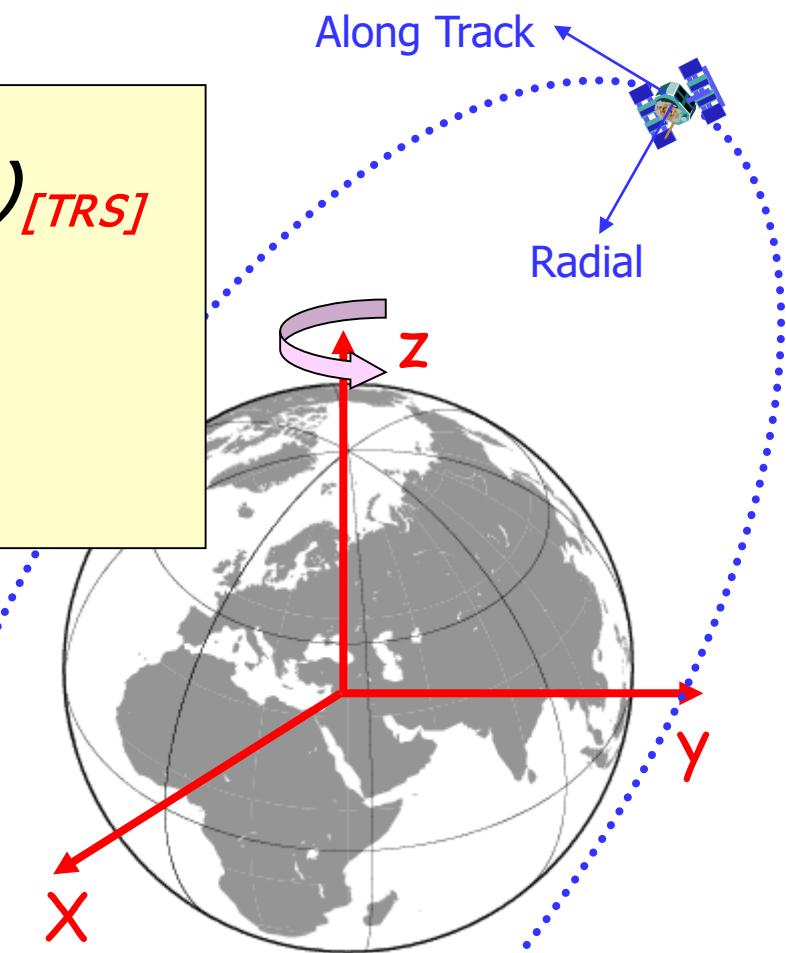
$$\begin{bmatrix} X_k \\ Y_k \\ Z_k \end{bmatrix} = \mathbf{R}_3(-\Omega_k) \mathbf{R}_1(-i_k) \mathbf{R}_3(-u_k) \begin{bmatrix} r_k \\ 0 \\ 0 \end{bmatrix}$$

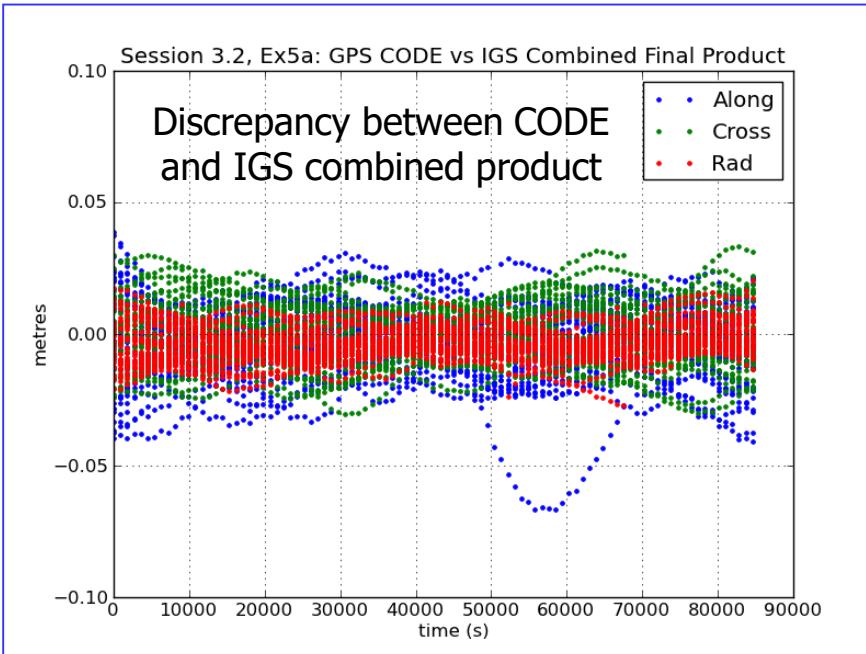
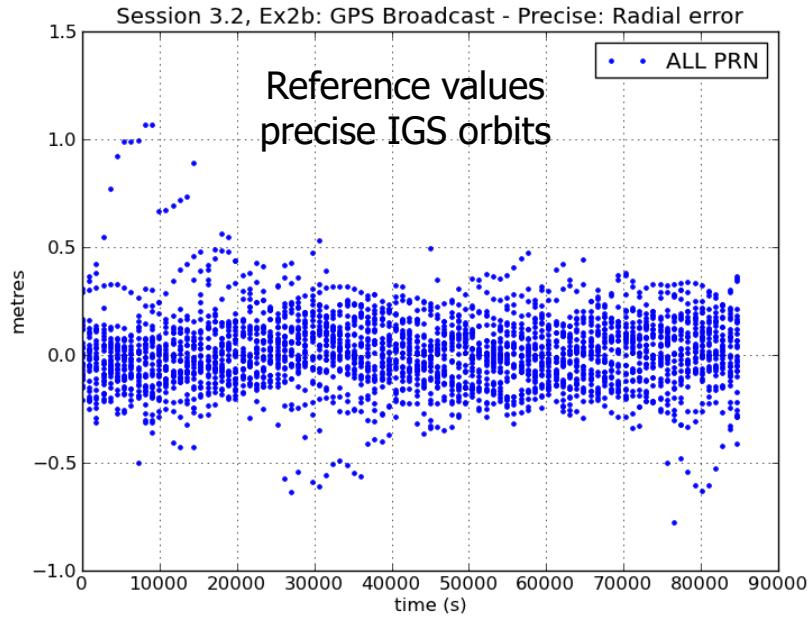
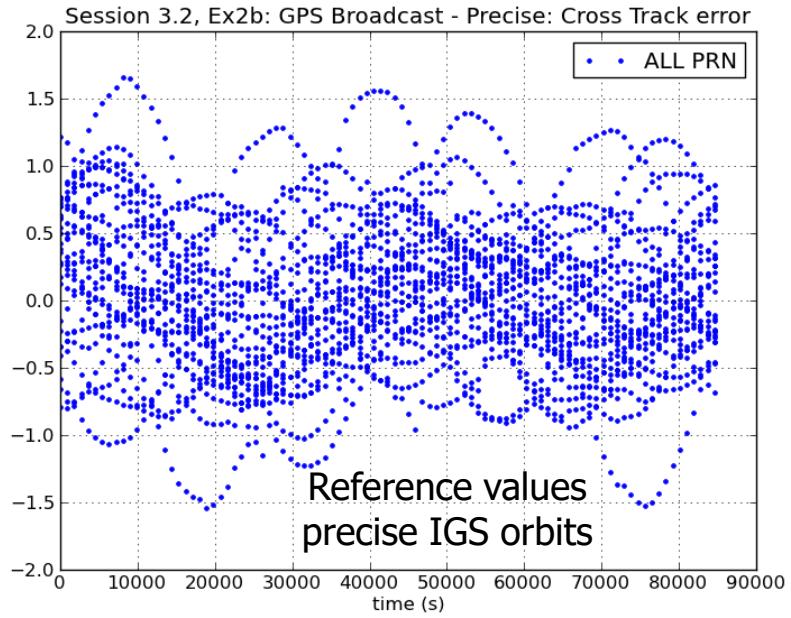
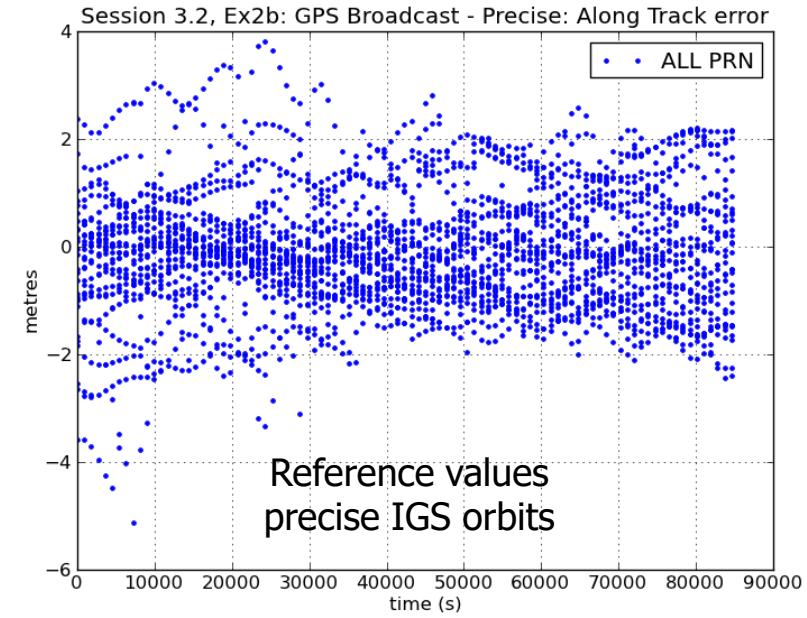


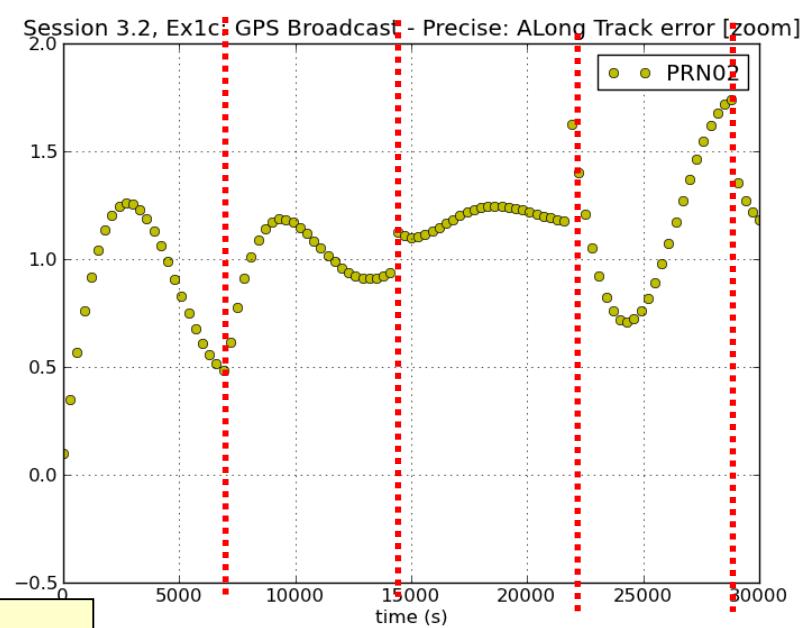
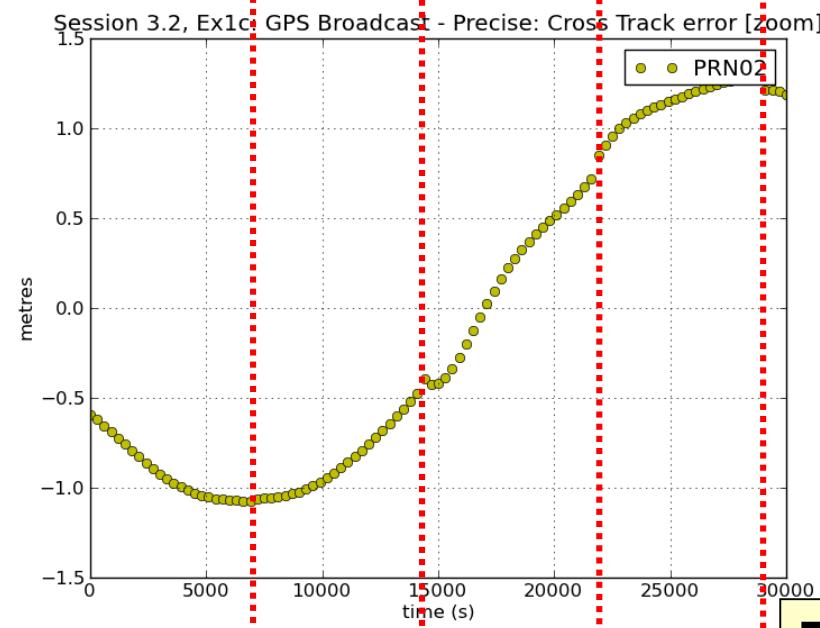
Conventional Terrestrial Reference System (**TRS**):

**Earth Centered, Earth-Fixed (ECEF) →**

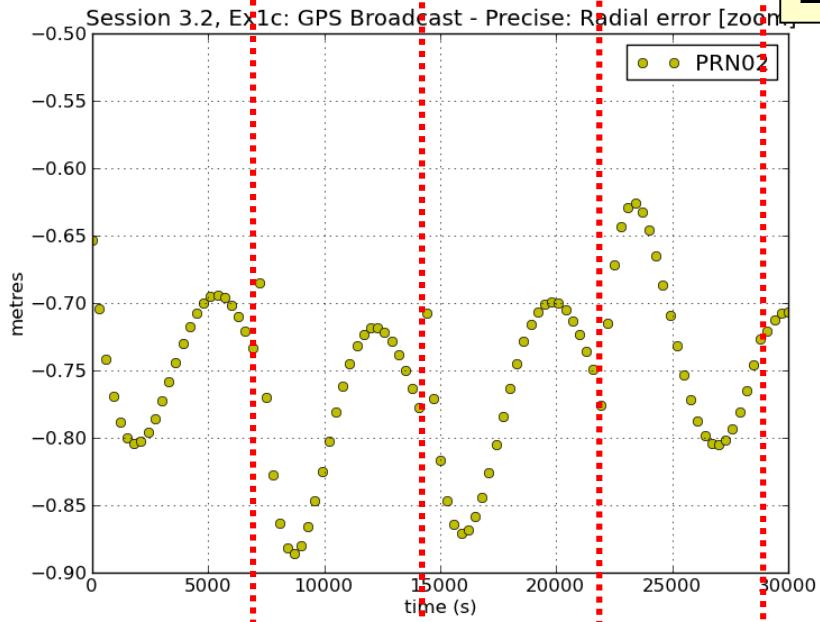
the reference system rotates with Earth.







**Zoom**



**Broadcast Orbit Updates**

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## 3.2 Computation of satellite coordinates from precise products.

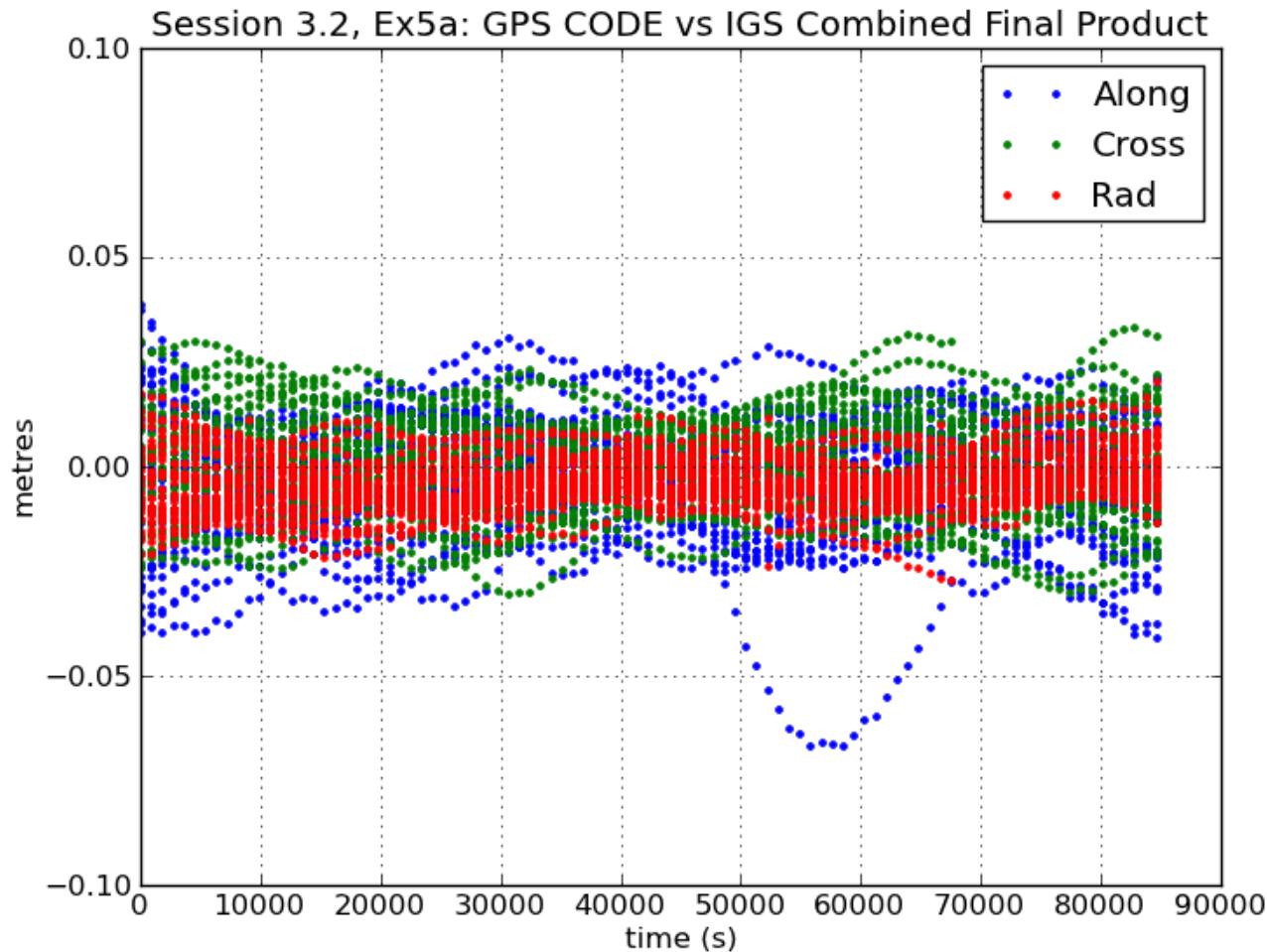
Precise orbits for GPS satellites can be found on the International GNSS Service (IGS) server <http://igscb.jpl.nasa.gov>

Orbits are given by  $(x,y,z)$  coordinates with a sampling rate of 15 minutes. The satellite coordinates between epochs can be computed by polynomial interpolation. A 10th-order polynomial is enough for a centimetre level of accuracy with 15 min data.

$$\begin{aligned} P_n(x) &= \sum_{i=1}^n y_i \frac{\prod_{j \neq i} (x - x_j)}{\prod_{j \neq i} (x_i - x_j)} \\ &= y_1 \frac{x - x_2}{x_1 - x_2} \dots \frac{x - x_n}{x_1 - x_n} + \dots \\ &\quad + y_i \frac{x - x_1}{x_i - x_1} \dots \frac{x - x_{i-1}}{x_i - x_{i-1}} \frac{x - x_{i+1}}{x_i - x_{i+1}} \dots \frac{x - x_n}{x_i - x_n} + \dots \\ &\quad + y_n \frac{x - x_1}{x_n - x_1} \dots \frac{x - x_{n-1}}{x_n - x_{n-1}} \end{aligned}$$

# IGS orbit and clock products (for PPP):

Discrepancy between CODE and IGS combined product.

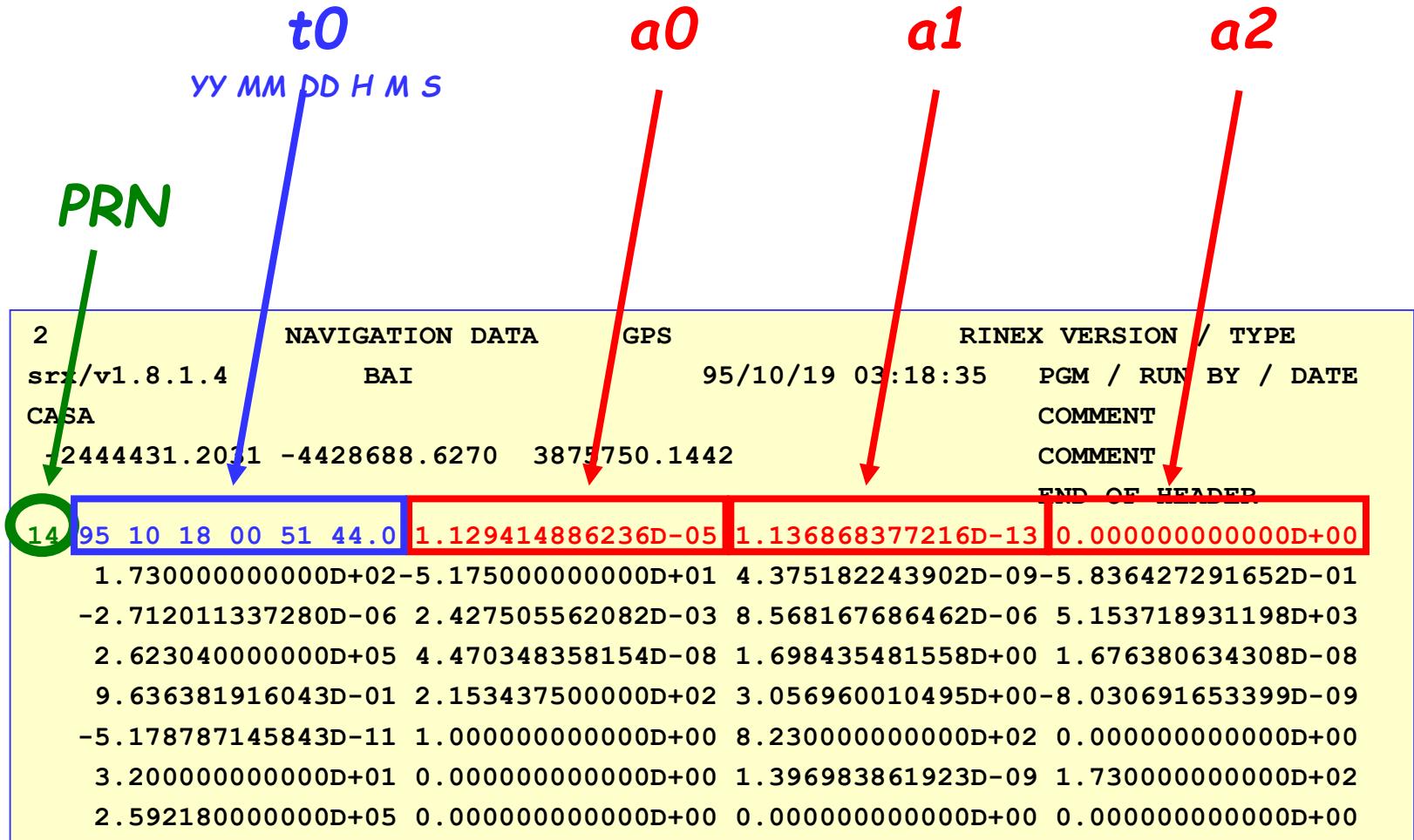


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4. GPS Satellite clock computation and accuracy
  - 4.1. From Broadcast Navigation Message.
  - 4.2. From precise products.

# GPS Satellite Clock computation: Broadcast message

$$dt^{sat} = a_0 + a_1(t - t_0) + a_2(t - t_0)^2$$



# Computation of satellite clocks from precise products

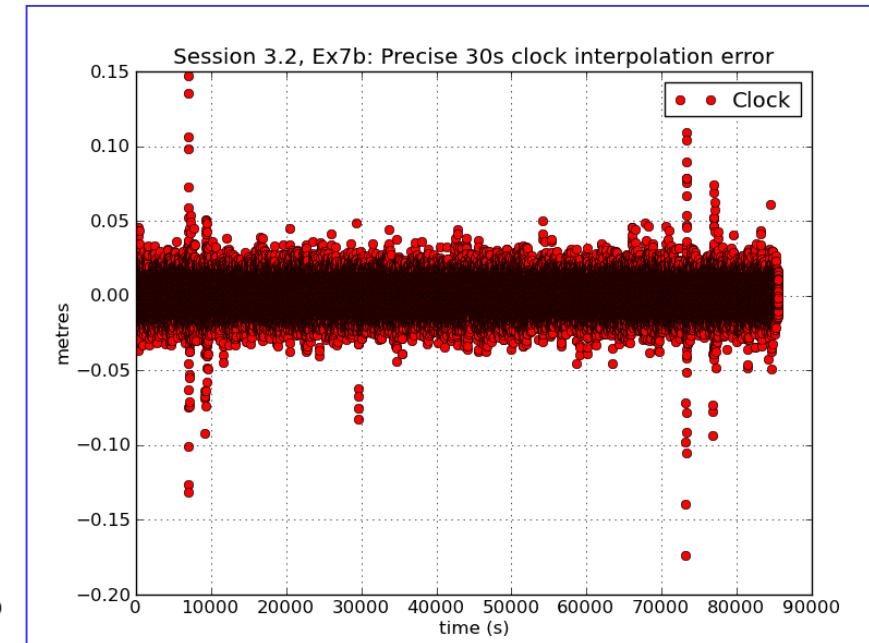
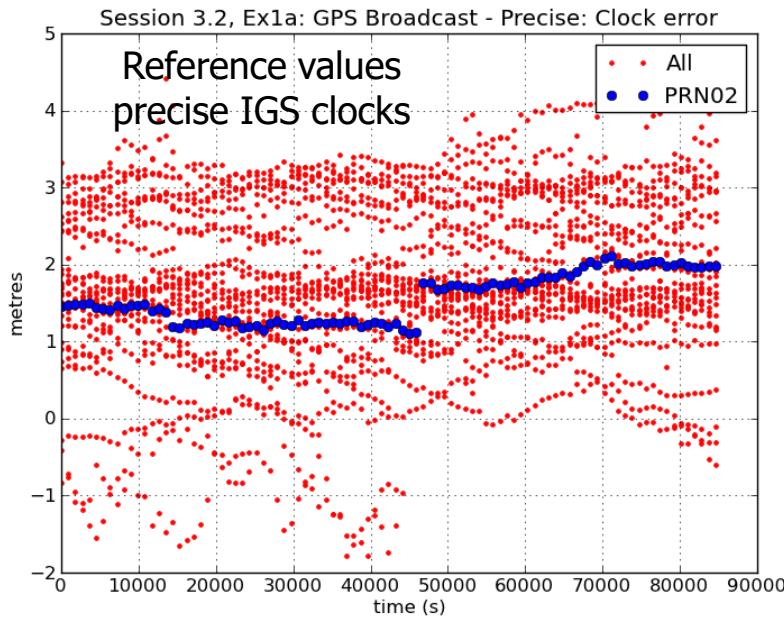
Precise clocks for GPS satellites can be found on the International GNSS Service (IGS) server <http://igscb.jpl.nasa.gov>

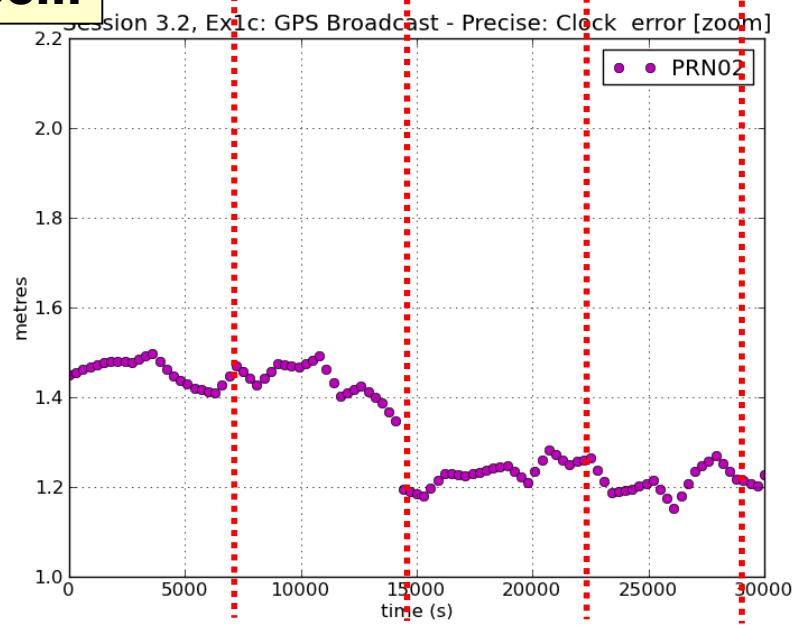
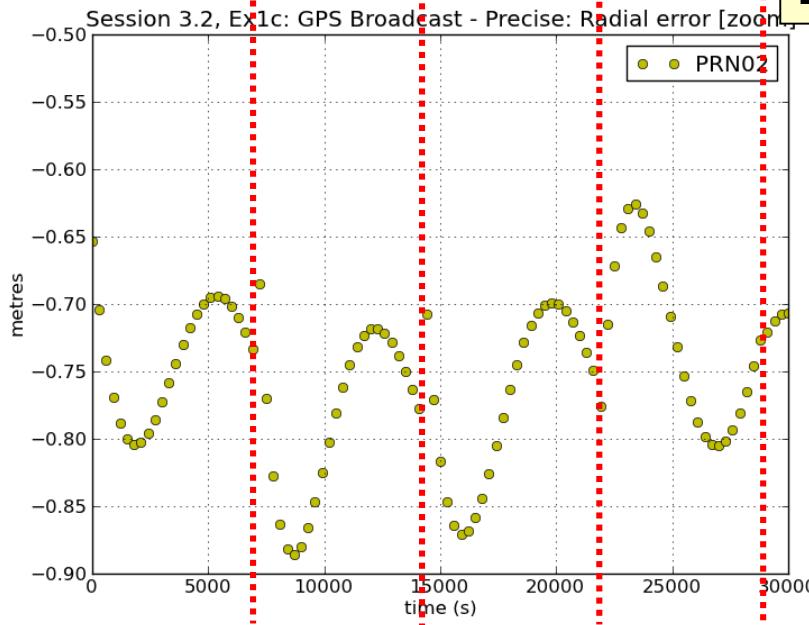
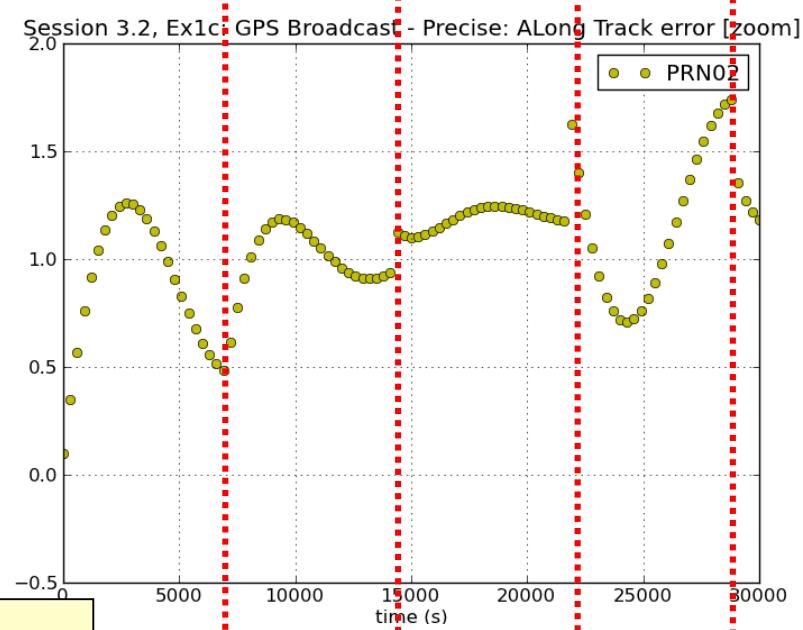
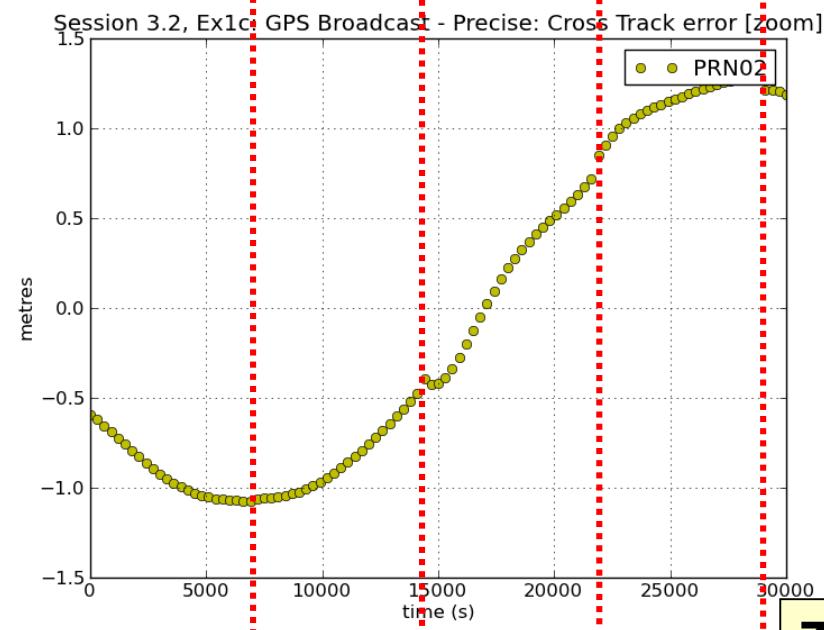
They are providing precise orbits and clock files with a sampling rate of 15 min (SP3 files), as well as precise clock files with a sample rate of 5 min and 30 s (CLK files).

Some centres also provide GPS satellite clocks with a 5 s sampling rate, like the ones obtained from the Crustal Dynamics Data Information System (CDDIS) site.

Stable clocks with a sampling rate of 30 s or higher can be interpolated with a first-order polynomial to a few centimetres of accuracy. Clocks with a lower sampling rate should not be interpolated, because clocks evolve as random walk processes.

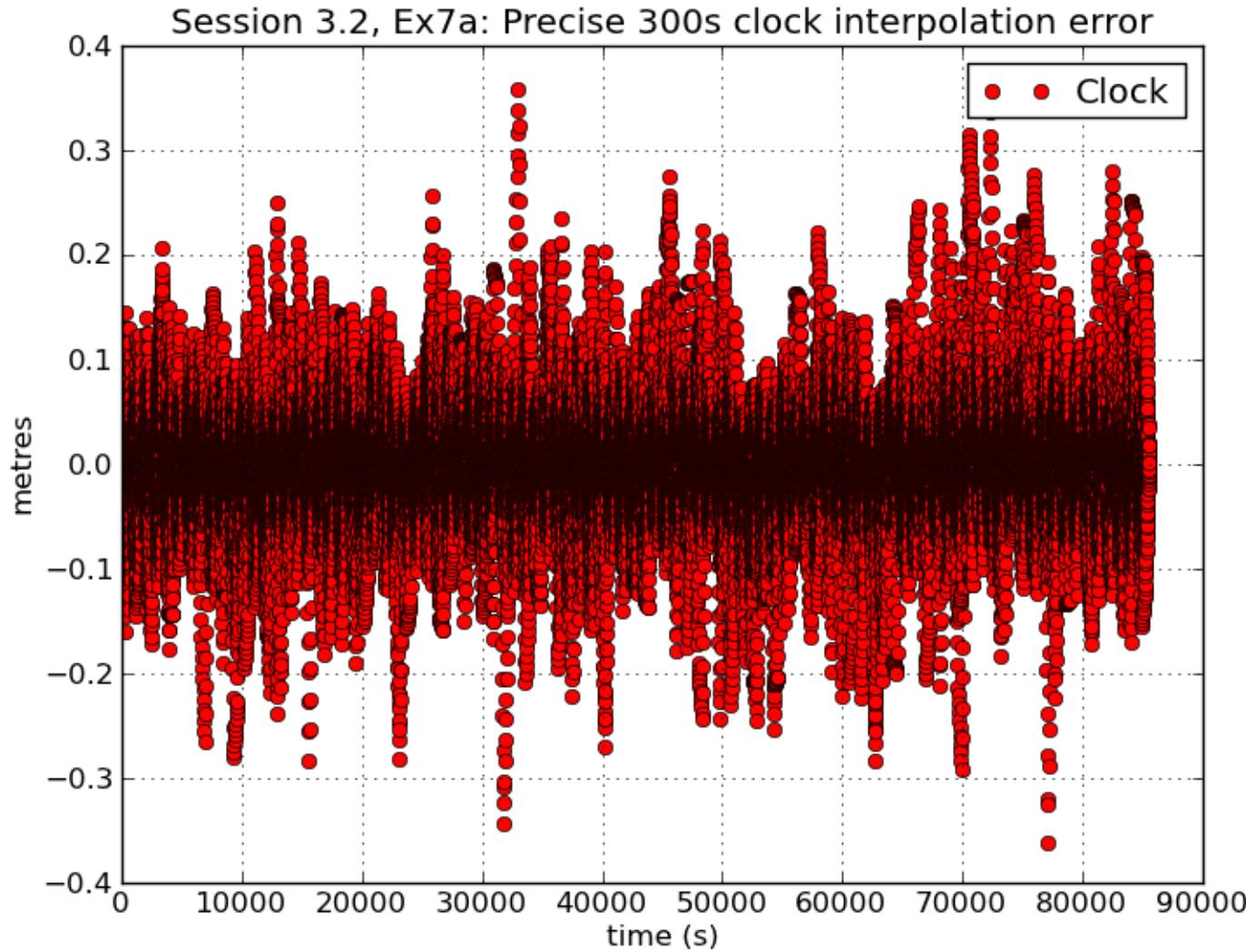
**SA=off**



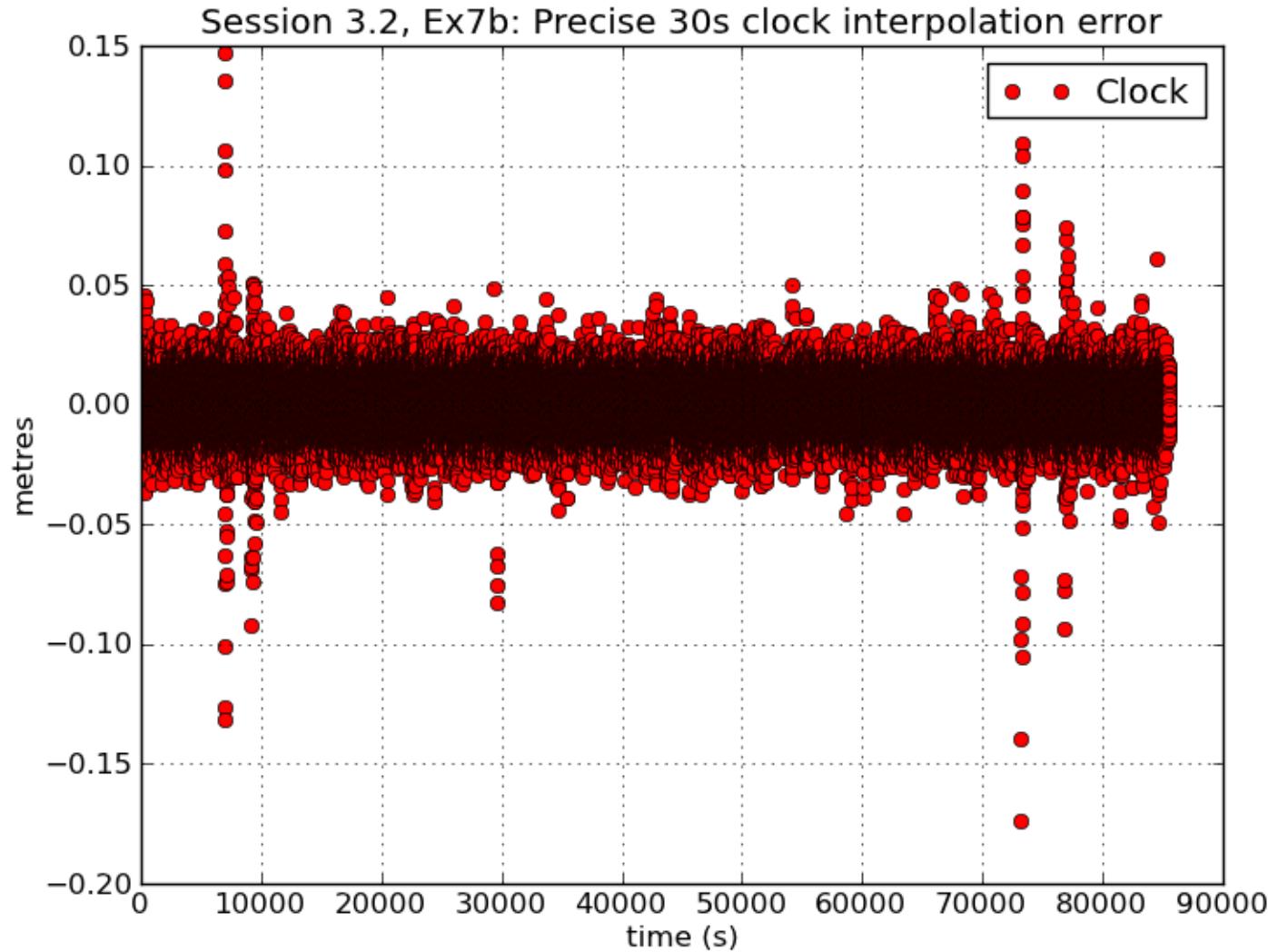


**Zoom**

## Precise Clock Interpolation: 300s samples

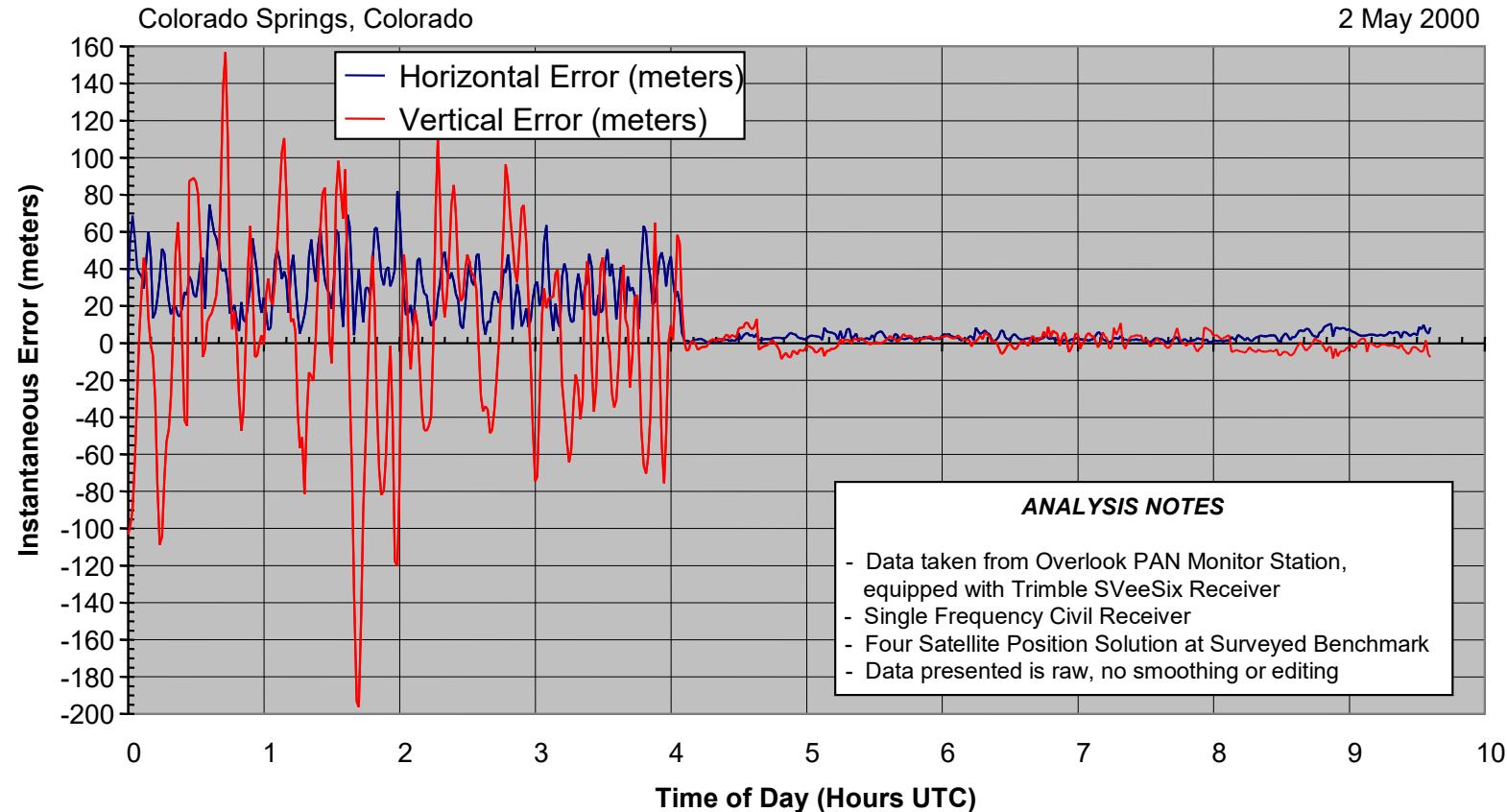


## Precise Clock Interpolation: 30s samples

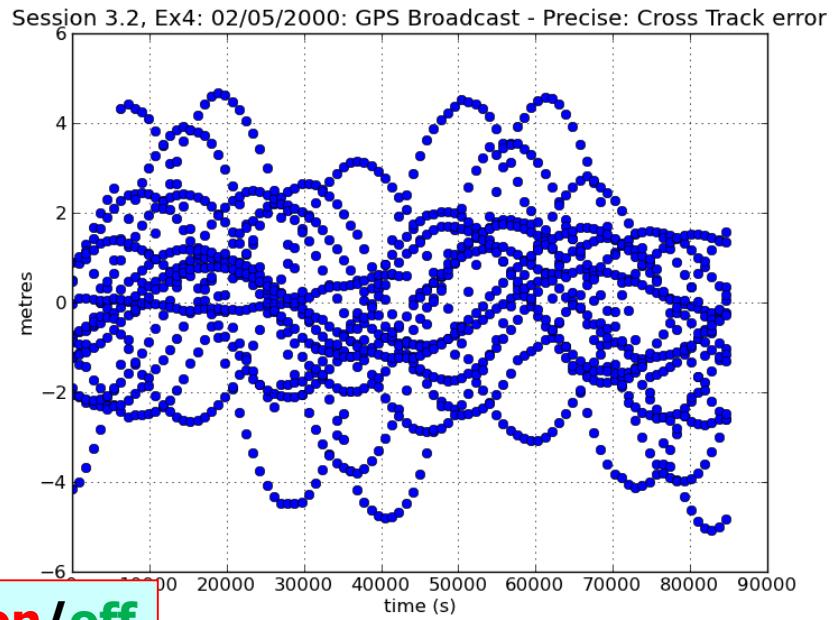
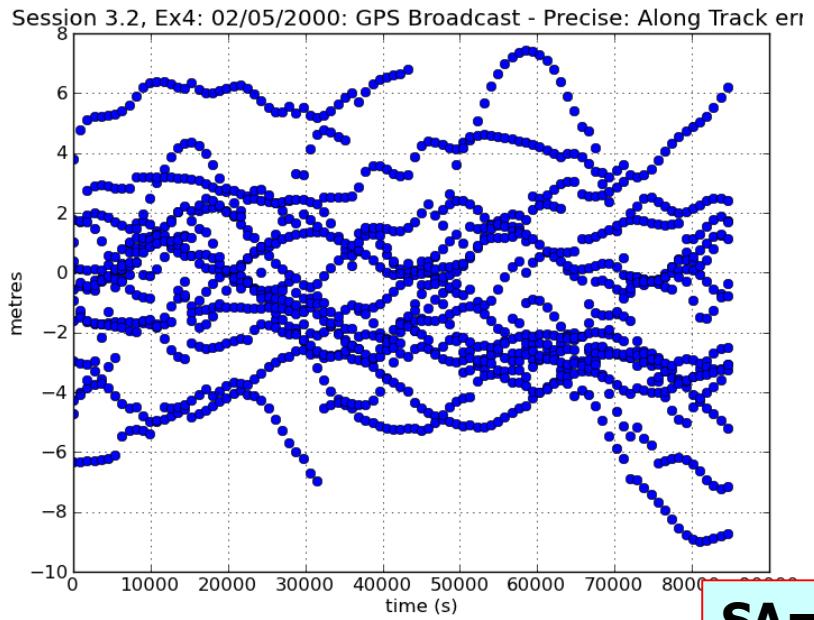


**Selective Availability (S/A):** Intentional degradation of satellite clocks and broadcast ephemeris. (from 25 March, 1990)

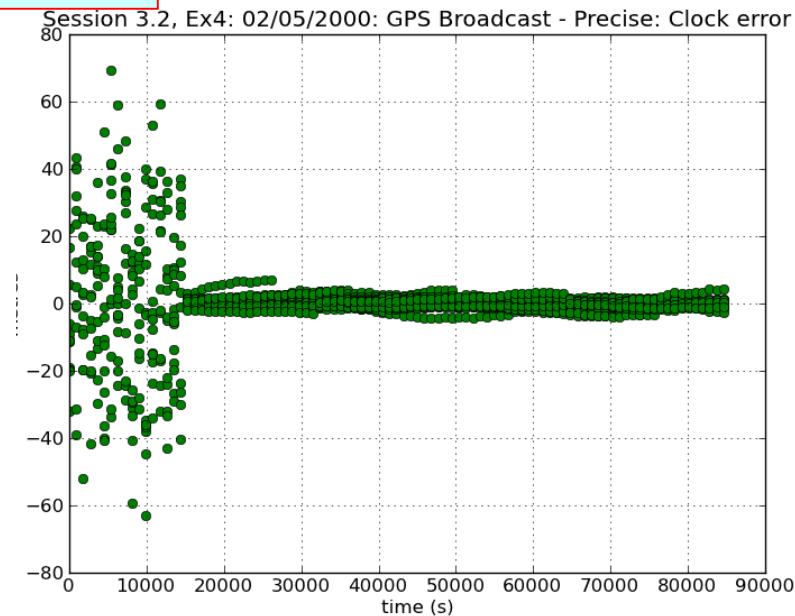
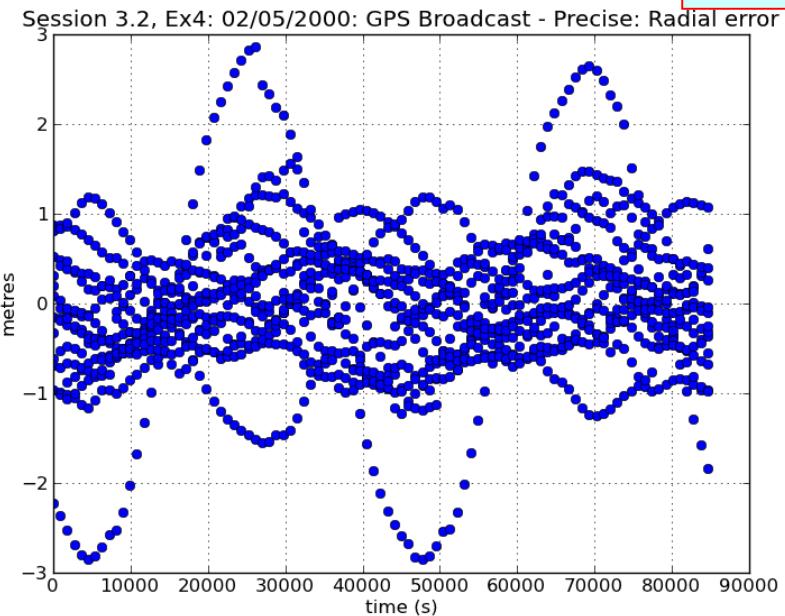
## GPS Before and After S/A was switched off

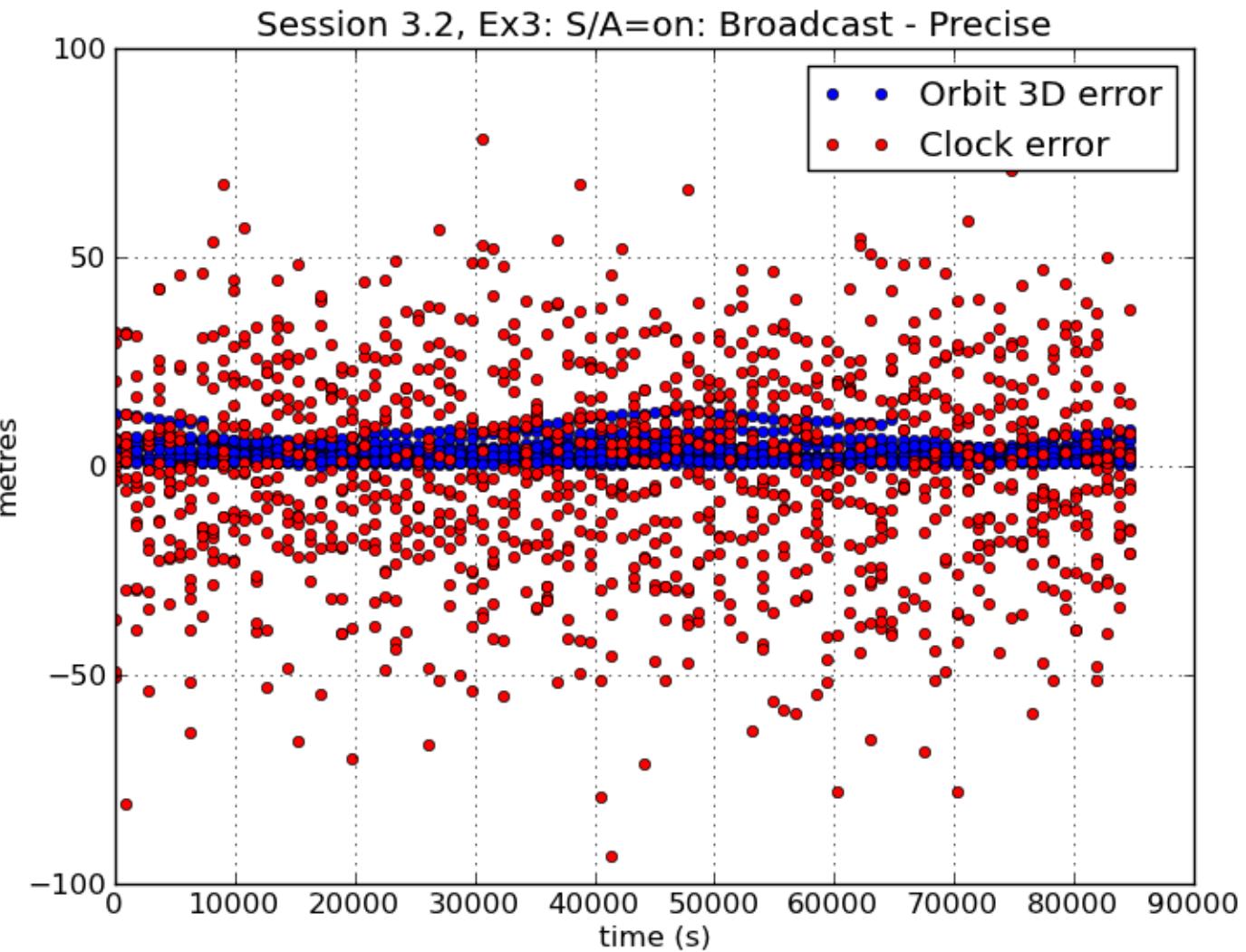


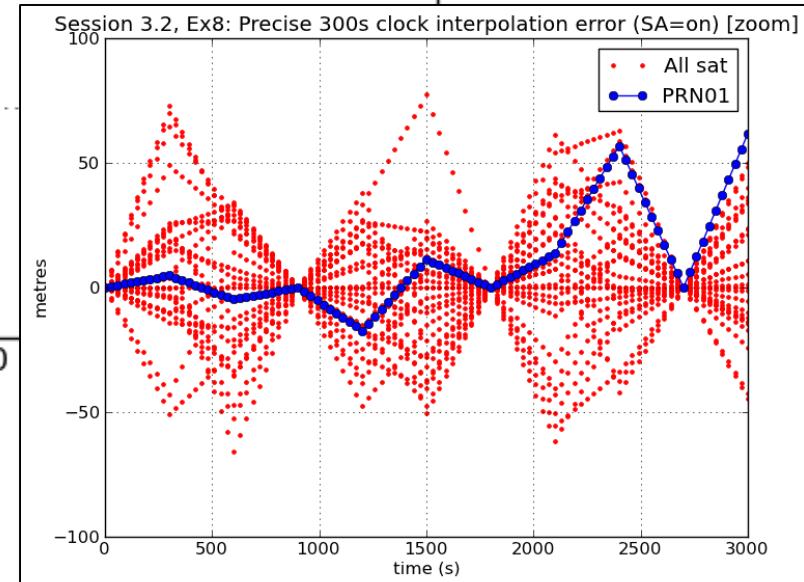
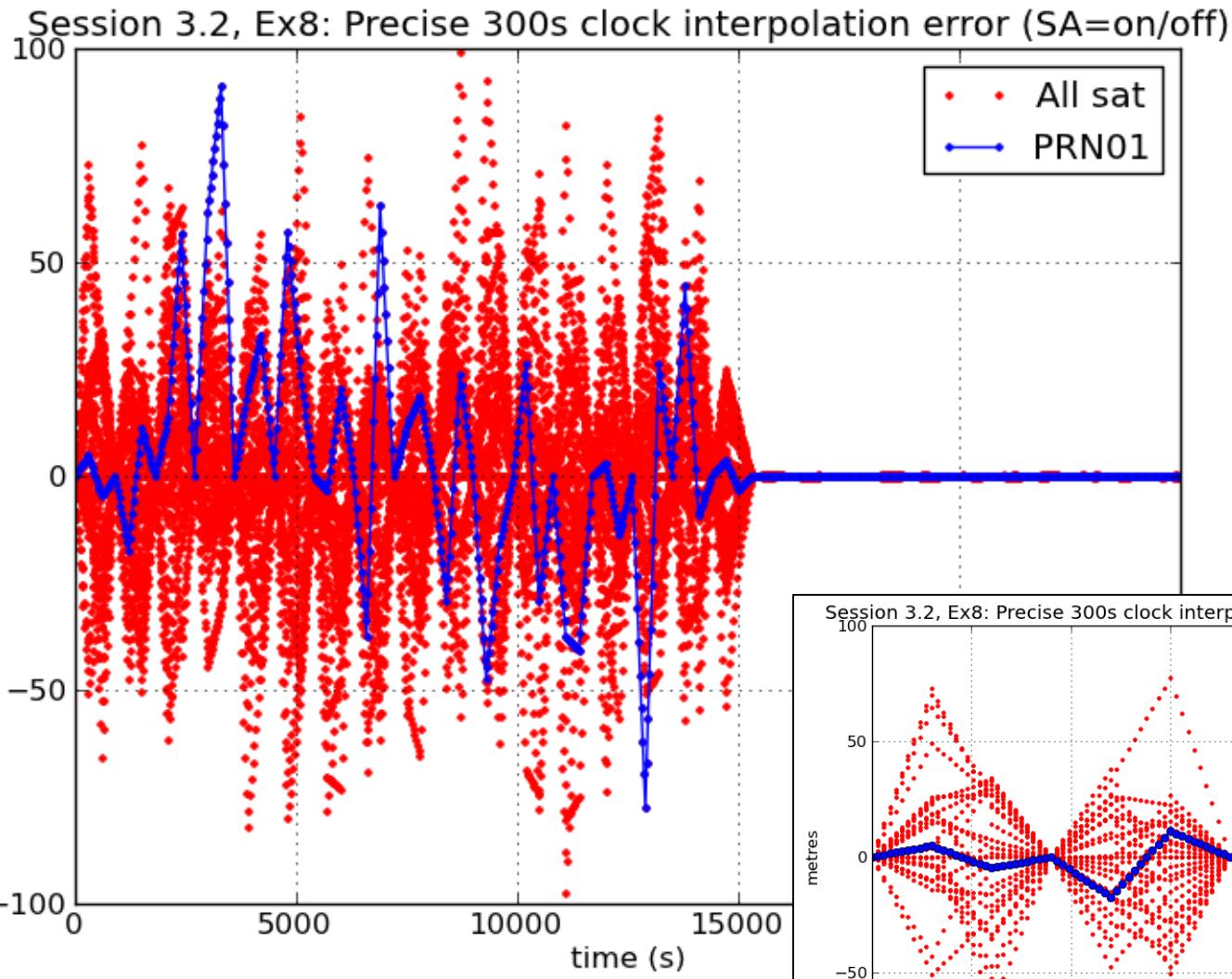
S/A was switched off at 2nd May 2000 and **Permanently removed in 2008**



**SA=on/off**







# References

- [RD-1] J. Sanz Subirana, J.M. Juan Zornoza, M. Hernández-Pajares, GNSS Data processing. Volume 1: Fundamentals and Algorithms. ESA TM-23/1. ESA Communications, 2013.
- [RD-2] J. Sanz Subirana, J.M. Juan Zornoza, M. Hernández-Pajares, GNSS Data processing. Volume 2: Laboratory Exercises. ESA TM-23/2. ESA Communications, 2013.
- [RD-3] Pratap Misra, Per Enge. Global Positioning System. Signals, Measurements, and Performance. Ganga –Jamuna Press, 2004.
- [RD-4] B. Hofmann-Wellenhof et al. GPS, Theory and Practice. Springer-Verlag. Wien, New York, 1994.

# Thank you

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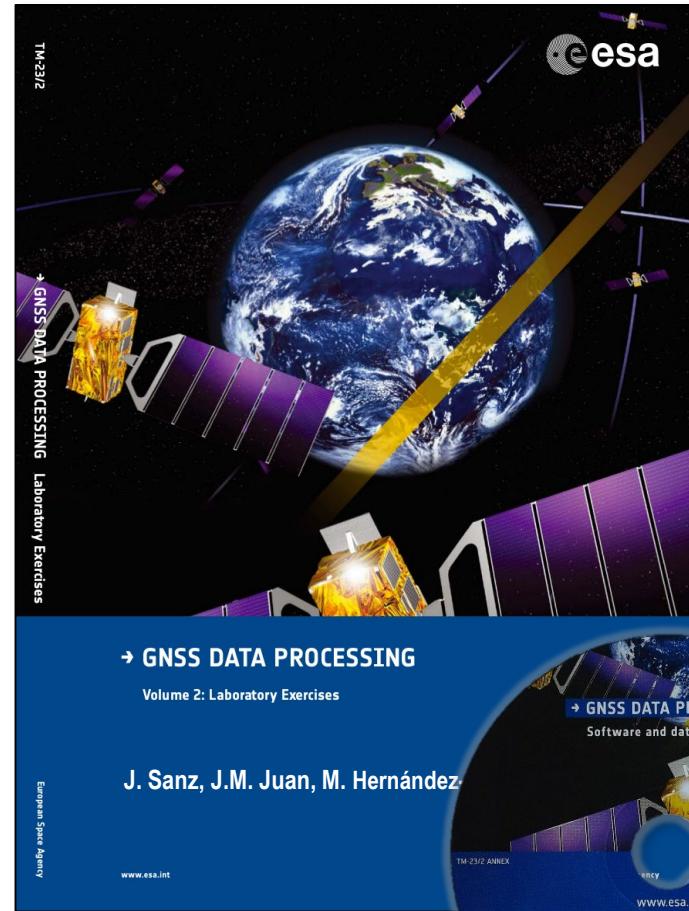
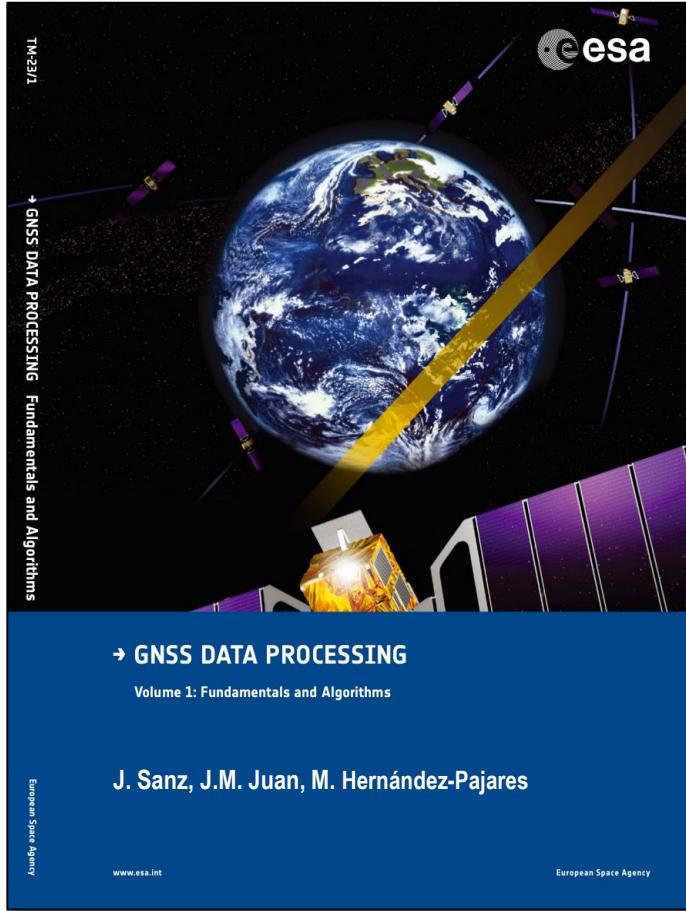
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This screenshot shows a web browser displaying a document titled 'GNSS\_Data\_processing\_course1.png' from the 'gAGE' website. The document is dated Tuesday, 09/03/2013, at 15:39, and is attributed to 'jaume.sanz'. The page content includes two main sections: 'Theory Slides' and 'Laboratory Slides', both associated with the 'GNSS Data Processing' course. Below these are images of a book titled 'GNSS DATA PROCESSING' and several presentation slides. A sidebar on the left lists various research areas and projects. On the right, there are sections for 'About us', 'Shortcuts', and a user login form. The overall layout is clean and professional, typical of an academic research group's website.



## **GNSS Data Processing, Vol. 1: Fundamentals and Algorithms. GNSS Data Processing, Vol. 2: Laboratory exercises.**